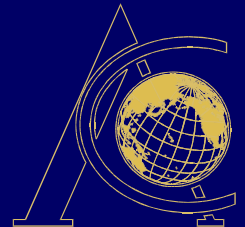


Imaging Basement & Intra-Sedimentary Horizons with Spectral Methods Applied to HRAM Data in Sirt Basin

Authors:

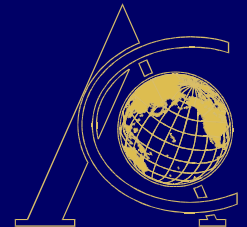
Irena Kivior, Sam Yates, Stephen Markham,
Francis Vaughan, Shiferaw Damte

EGM 2007 International Workshop
Capri, April 15-18 2007



Outline

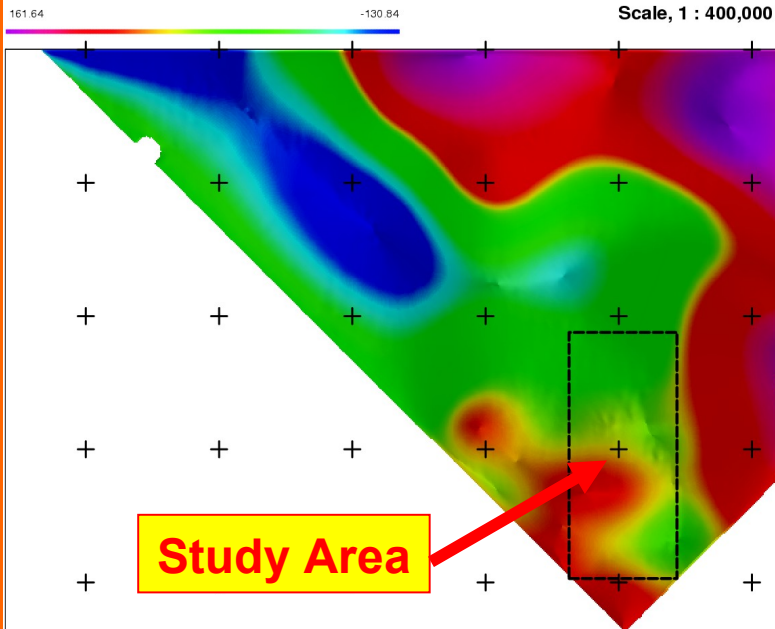
- **Introduction**
- **Methodology**
- **Horizon Interpretation**
 - Basement
 - Top Nubian
 - Unit 4 of Nubian SST
- **Comparison with Well and Seismic data**
- **Conclusions**



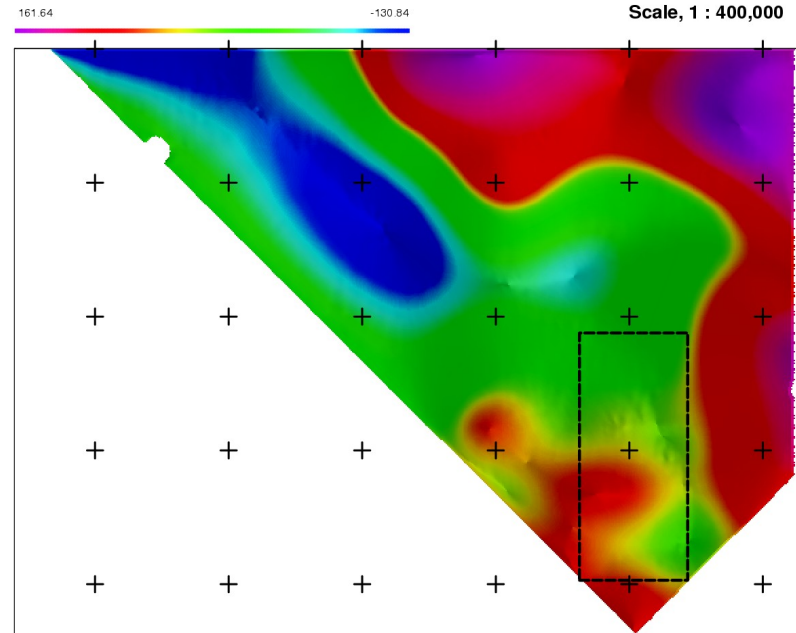
Introduction

- HRAM data was analysed over part of the Sirt Basin to image the Magnetic Horizons
- Survey Parameters:
 - Flight Line Spacing 500m
 - Tie Line Spacing 2500m
 - Flight Altitude 100m
 - Flight Line Direction NE-SW

Total Magnetic Intensity

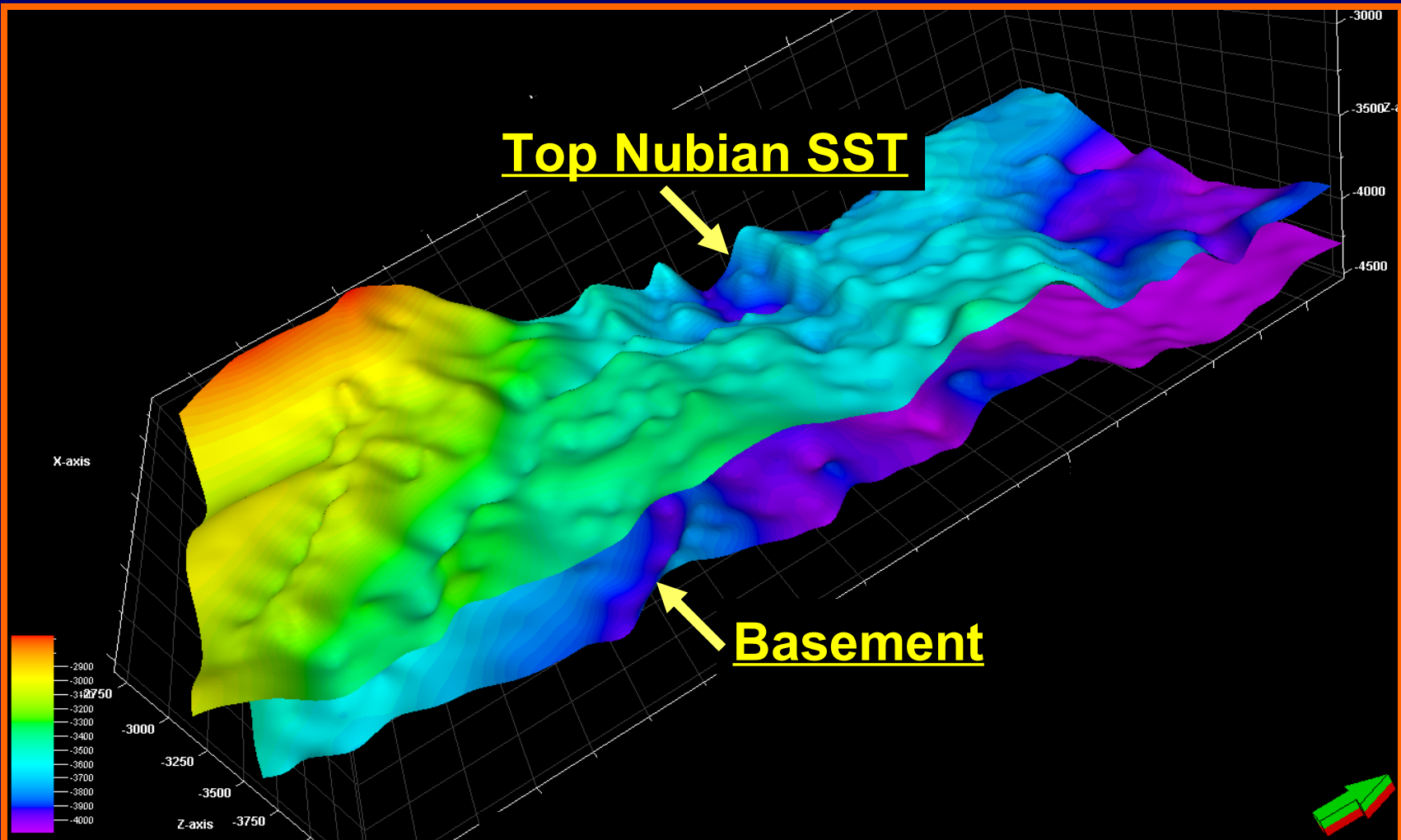


Reduced to Pole (RTP)



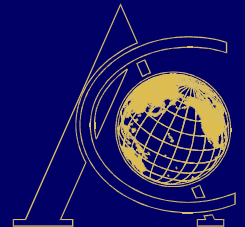
Introduction

- Spectral methods were used to map Basement & Top of Nubian SST Intra-Sedimentary Horizon.



Outline

- Introduction
- Methodology
- Horizon Interpretation
 - Basement
 - Top Nubian
 - Unit 4 of Nubian SST
- Comparison with Well and Seismic data
- Conclusions



Methodology

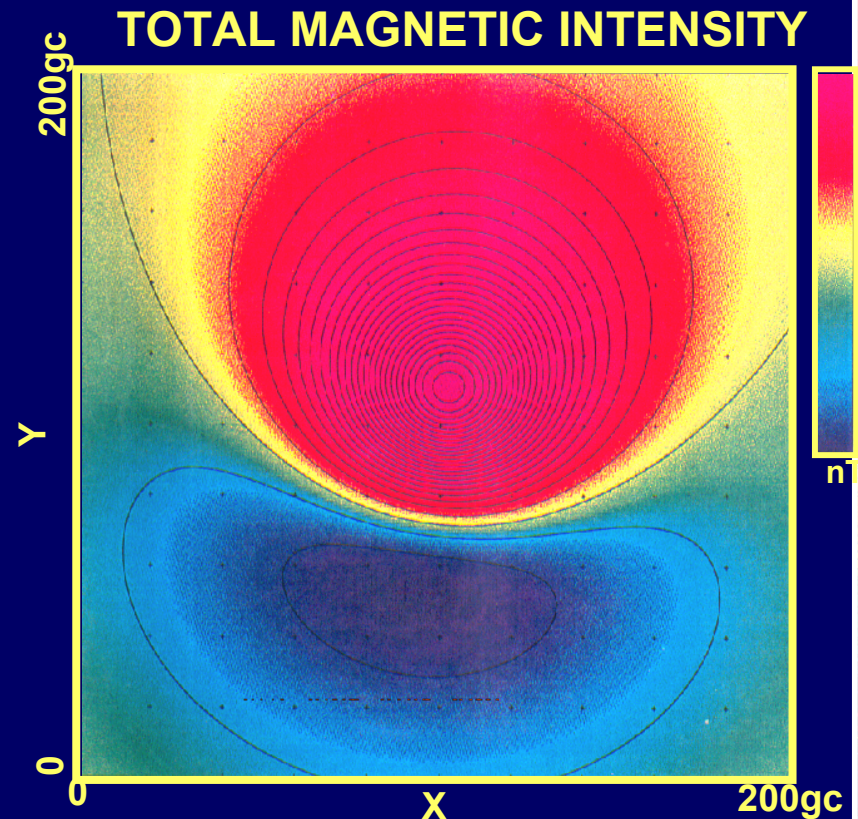
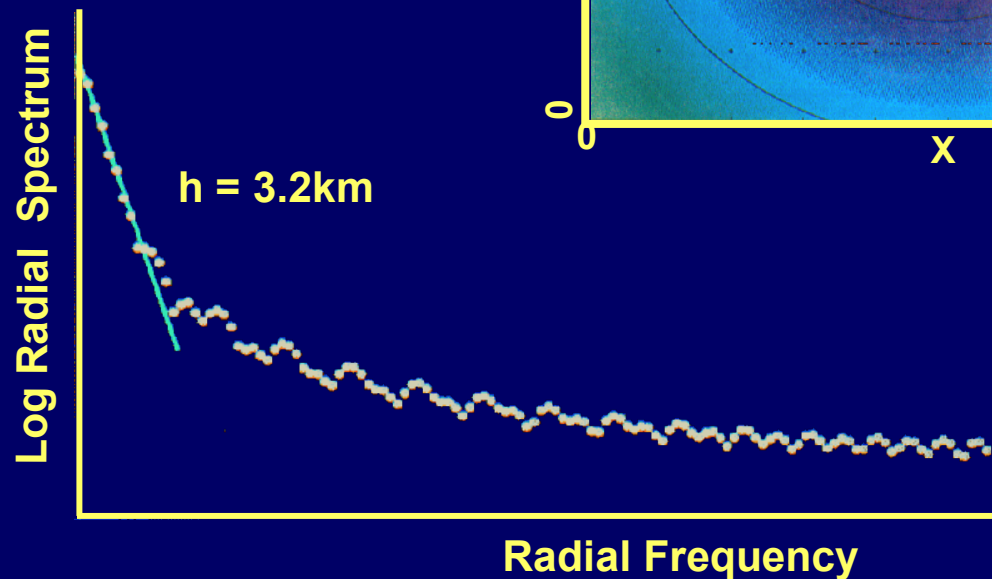
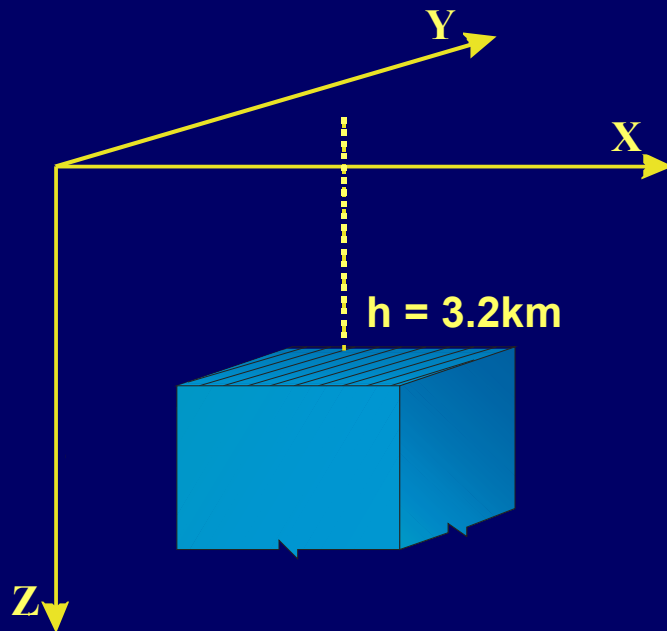
Energy Spectral Analysis

Basic Principles

- **Theoretical Models**
 - **Single Vertical Prism**
 - **Multi-prisms Model**



SINGLE PRISM MODEL

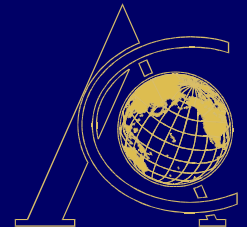


Methodology

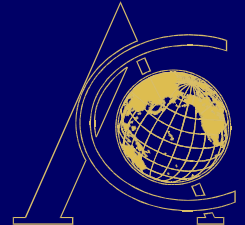
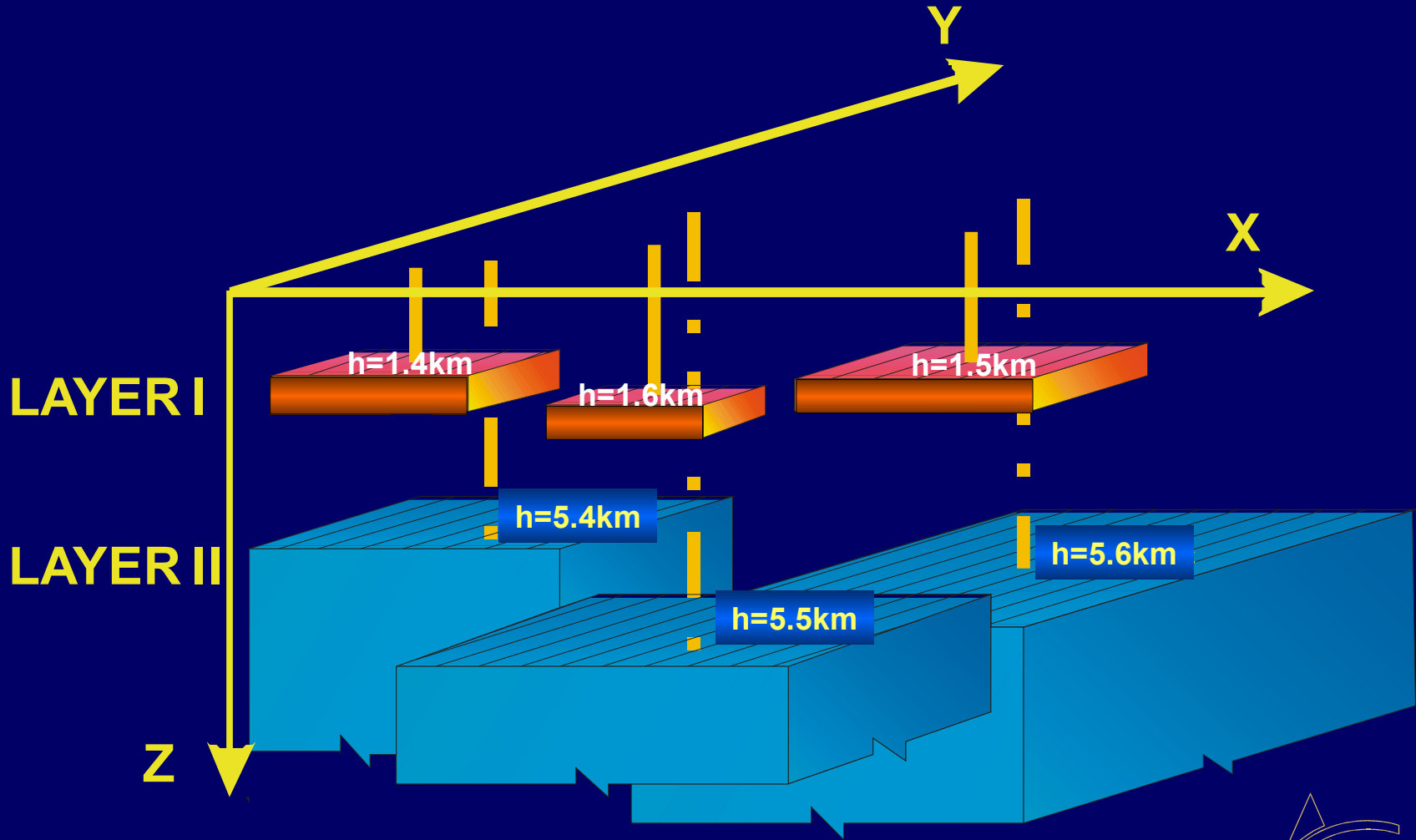
Energy Spectral Analysis

Basic Principles

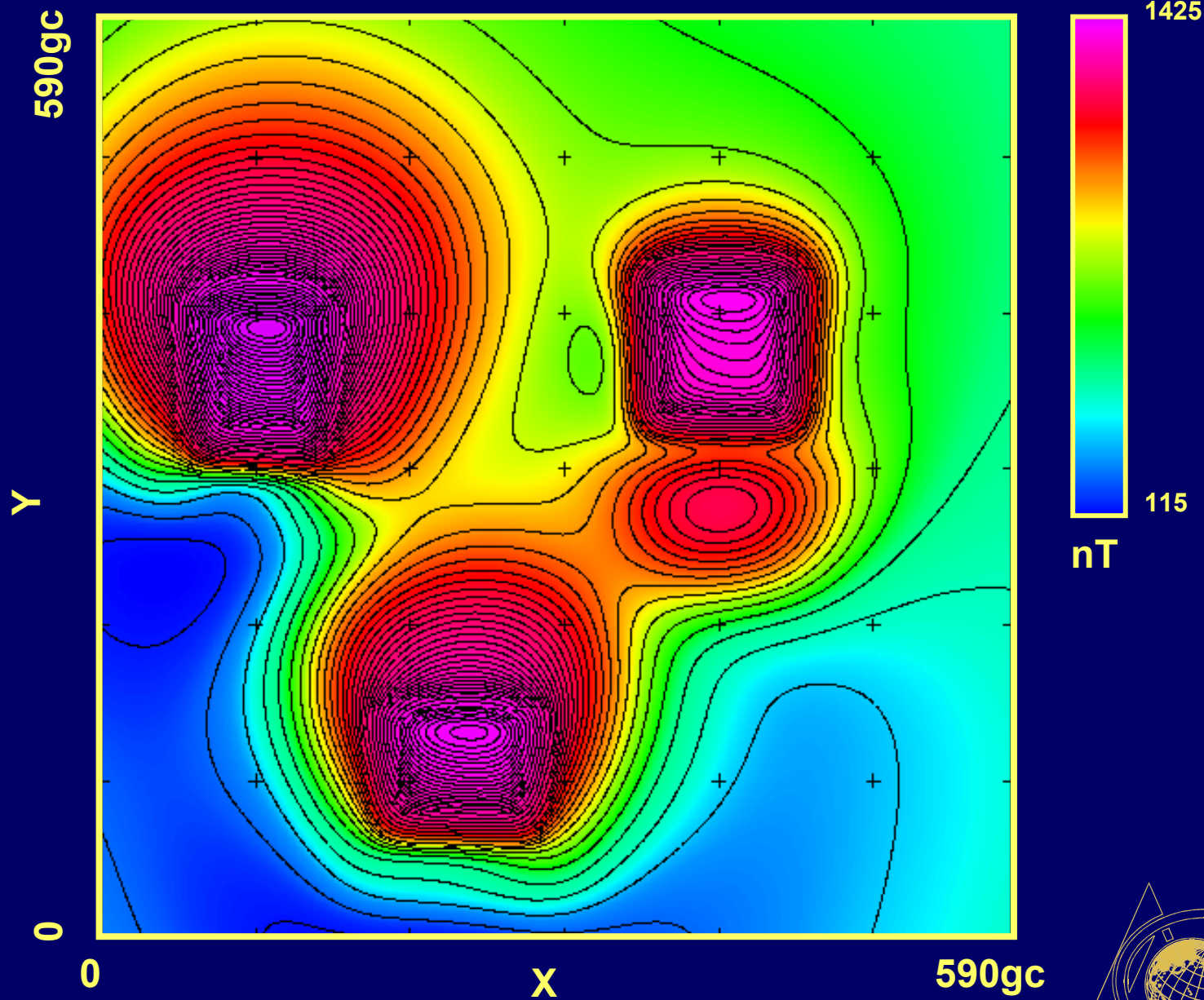
- **Theoretical Models**
 - Single Vertical Prism
 - Multi-prisms Model



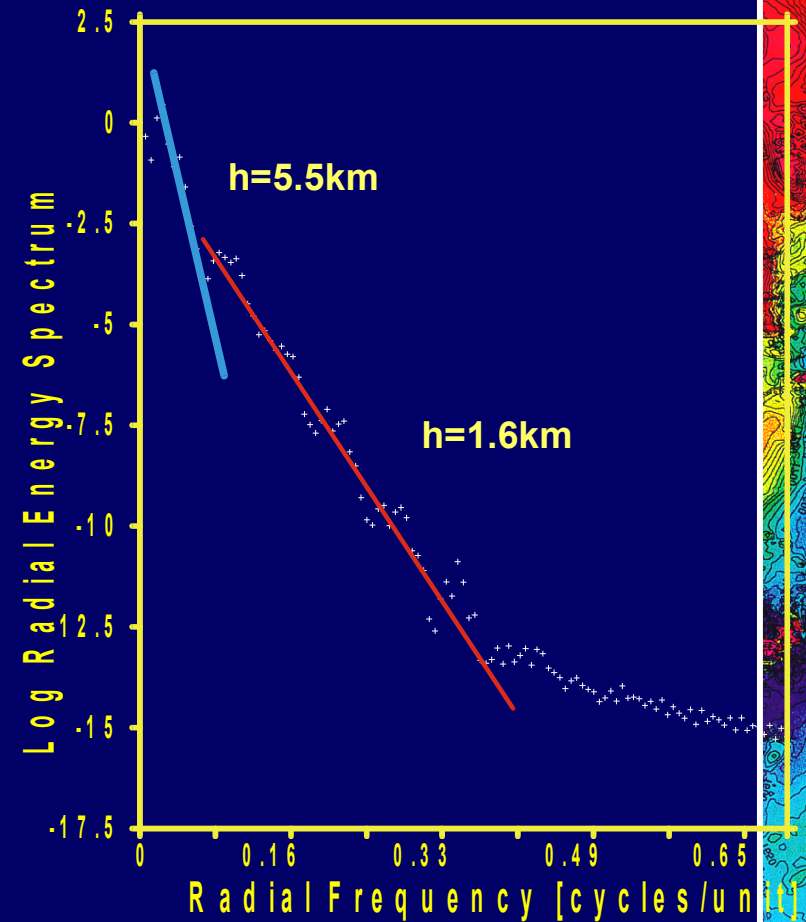
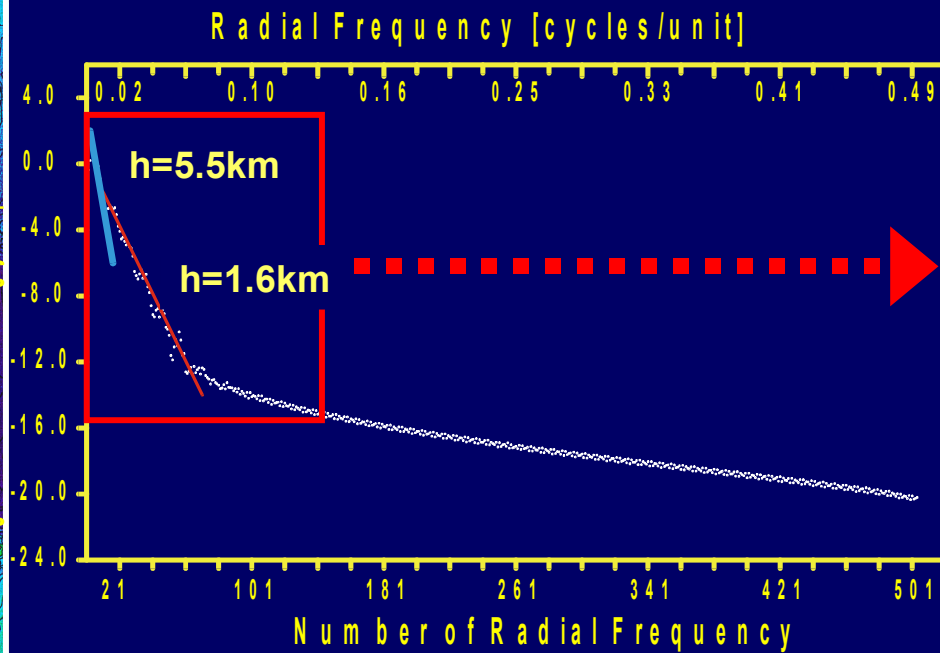
Methodology: Multi-Prism Model



MULTI-PRISMS MODEL - TMI



MULTI-PRISMS MODEL: ENERGY SPECTRUM

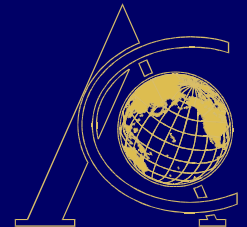
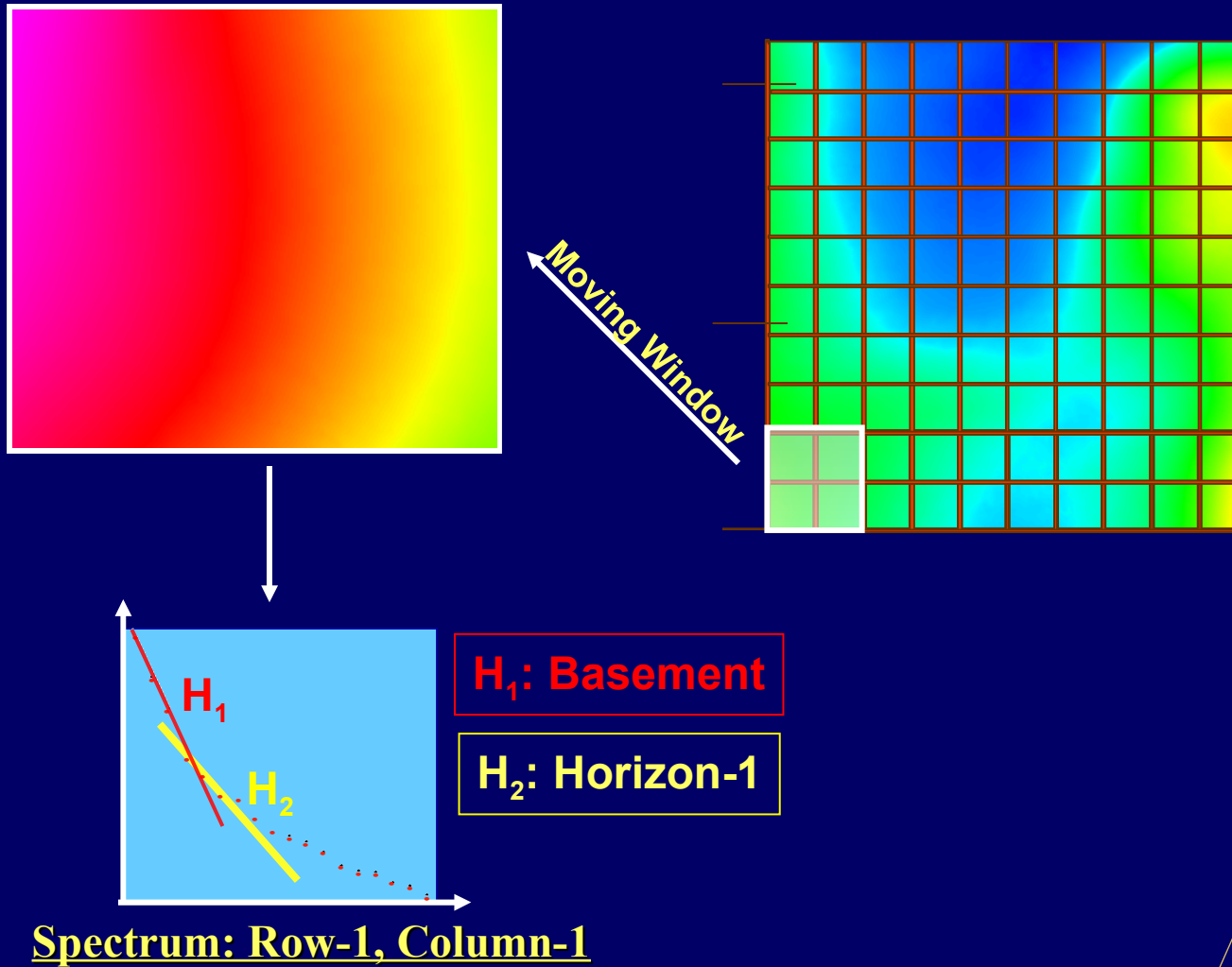


DETAIL OF LOW
FREQUENCIES



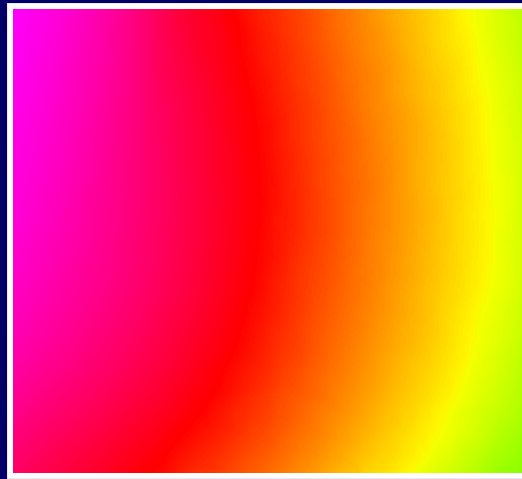
Methodology

ESA Moving Window

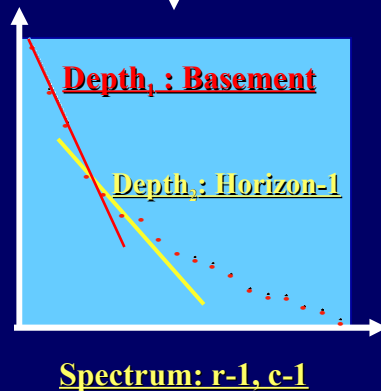
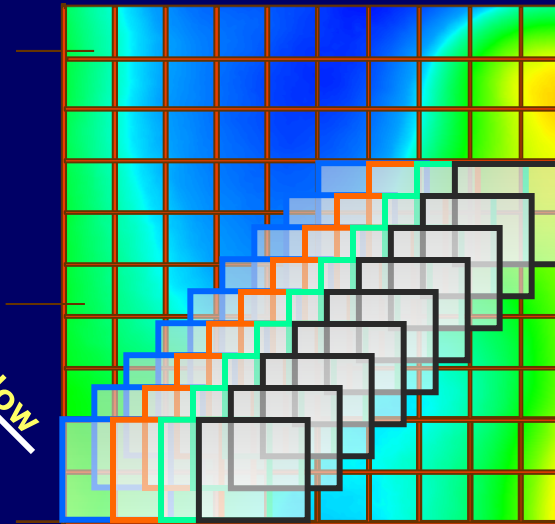


Methodology

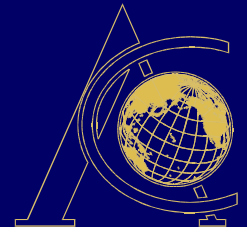
ESA Moving Window



Moving Window



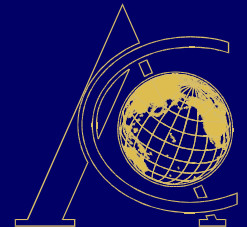
- Moving-Window increment e.g. 1x1gc
- Depth to Horizons every gc



Methodology: ESA-MW

Depth to Magnetic Horizons

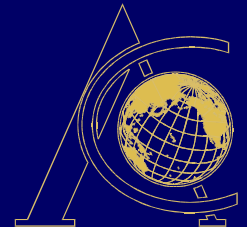
- A large number of spectra are computed for each interpreted horizon
- Each horizon is imaged using
 - Semi-automatic depth calculation
 - Fully-automatic depth calculation - AutoESA



Methodology: ESA-MW

Depth to Magnetic Horizons

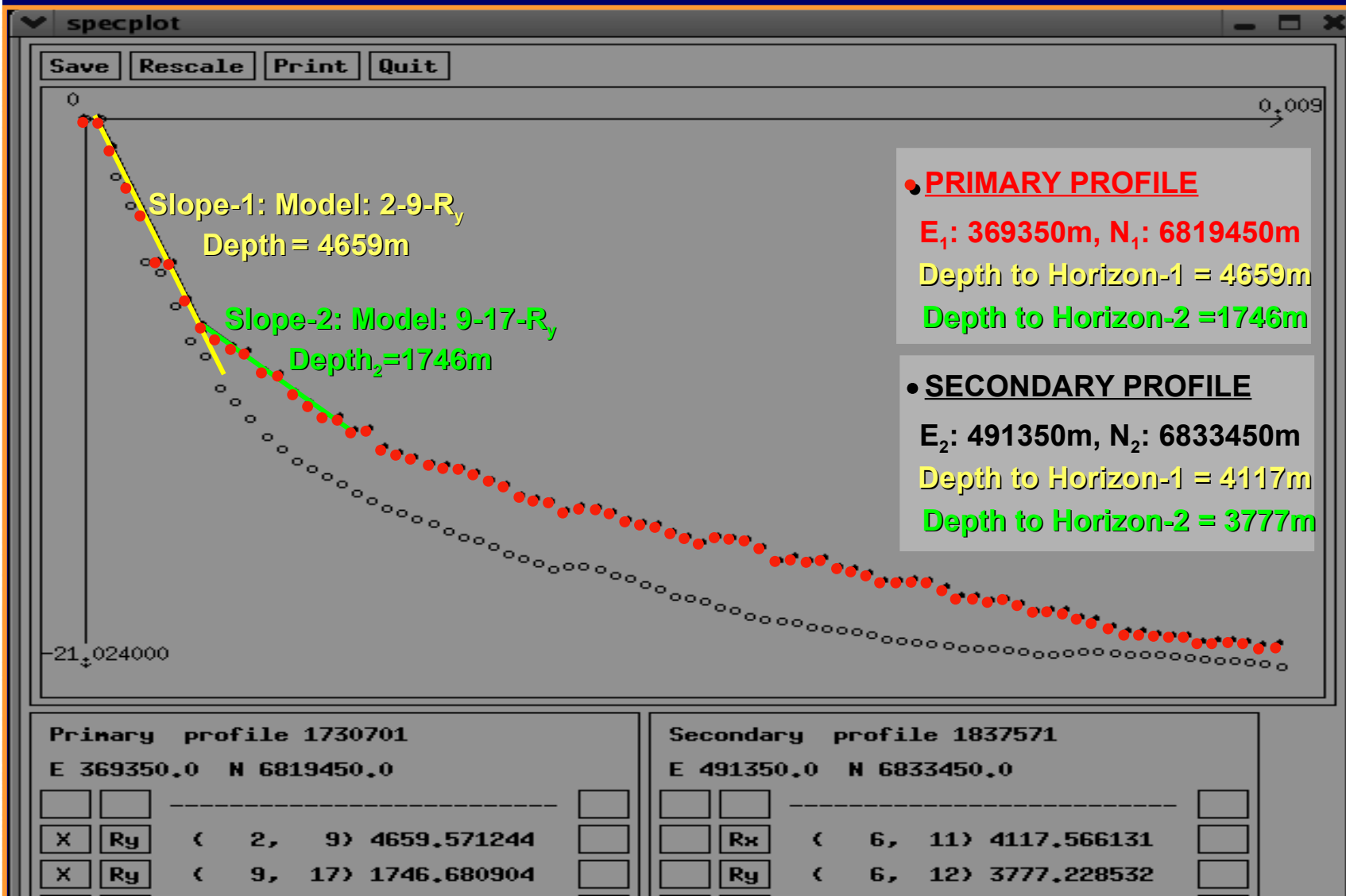
- **Semi-Automatic Depth Calculation**
- **AutoESA: Fully-automatic Depth Calculation**



Sedimentary Basin

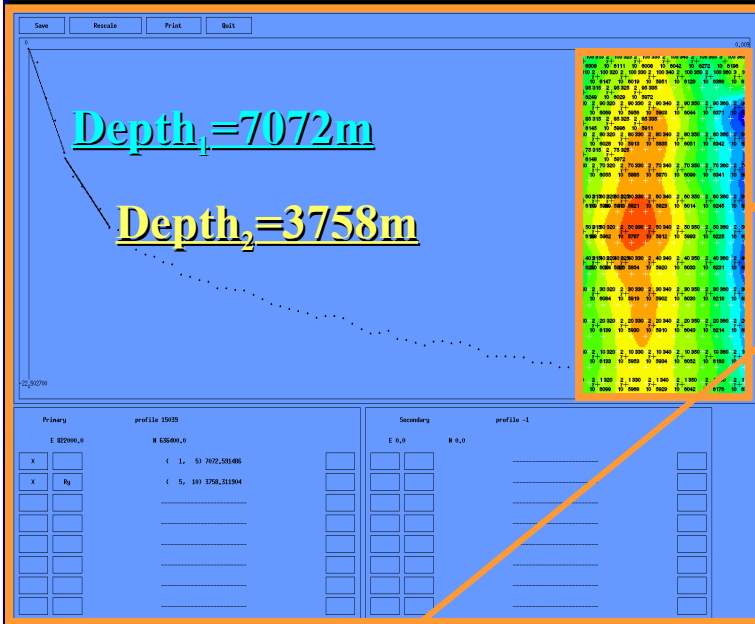
ESA-MW: Semi-Automatic Depth Interpretation

Spectra Window 12.8x12.8km: Horizon-1 & Horizon-2

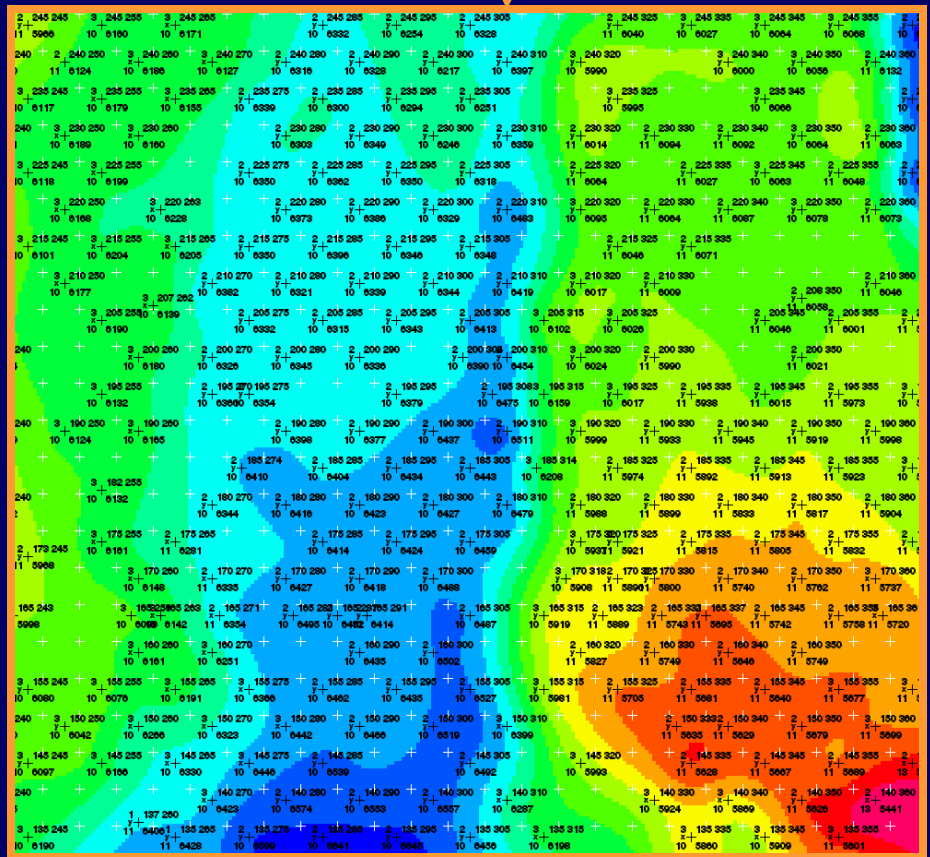
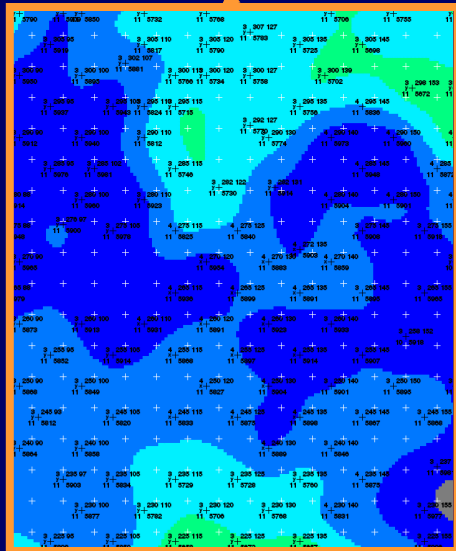


ESA 'Moving Window'

Semi-Automatic Depth Calculation

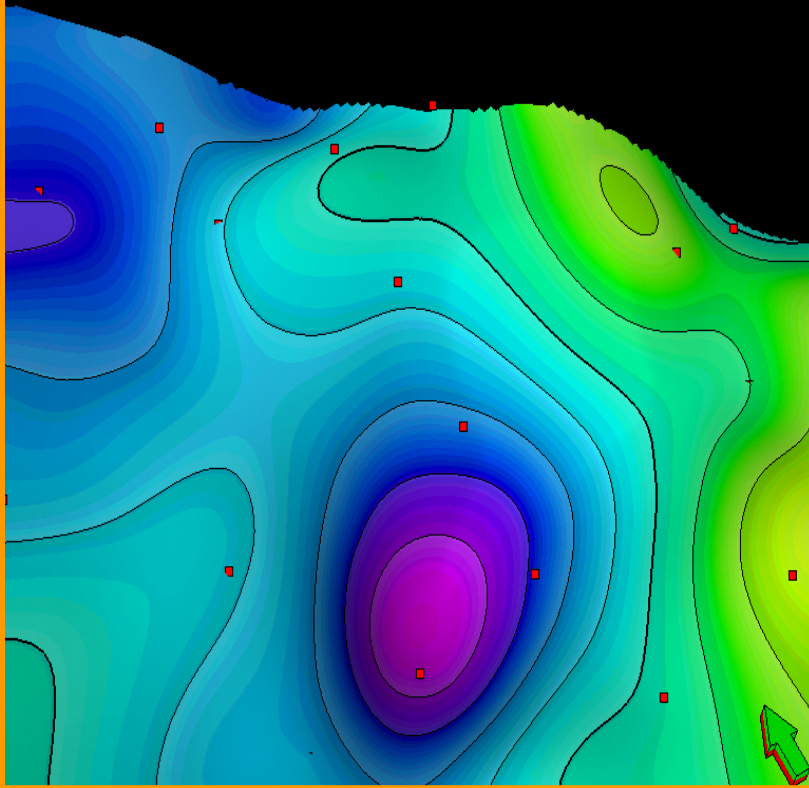


Live Image of Interpreted
Magnetic Basement

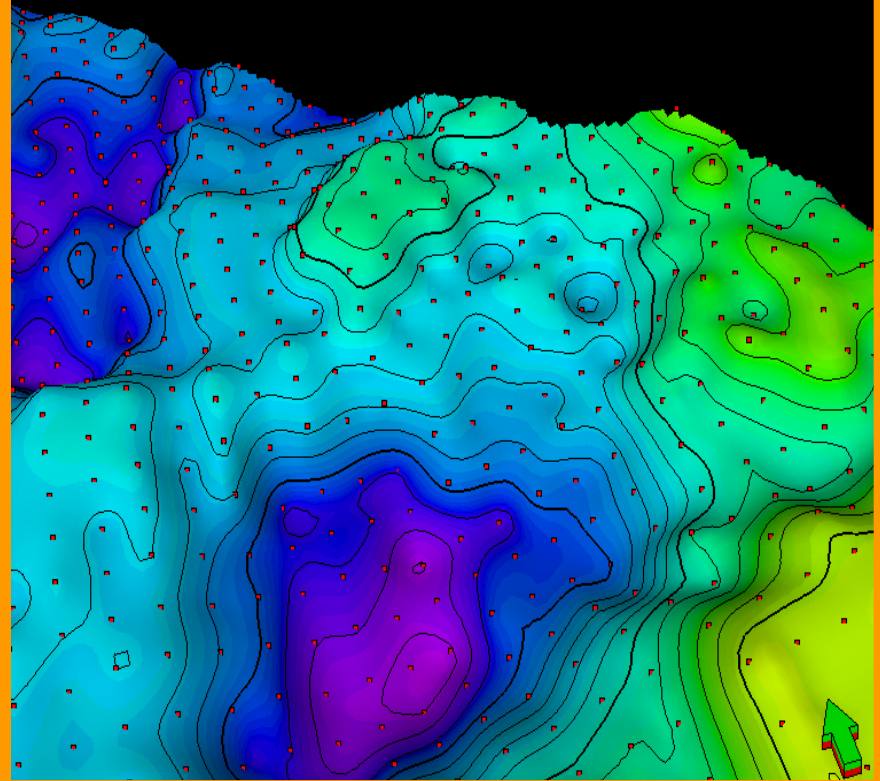


Low & High Resolution Interpretation of HRAM Data

Low Resolution Results

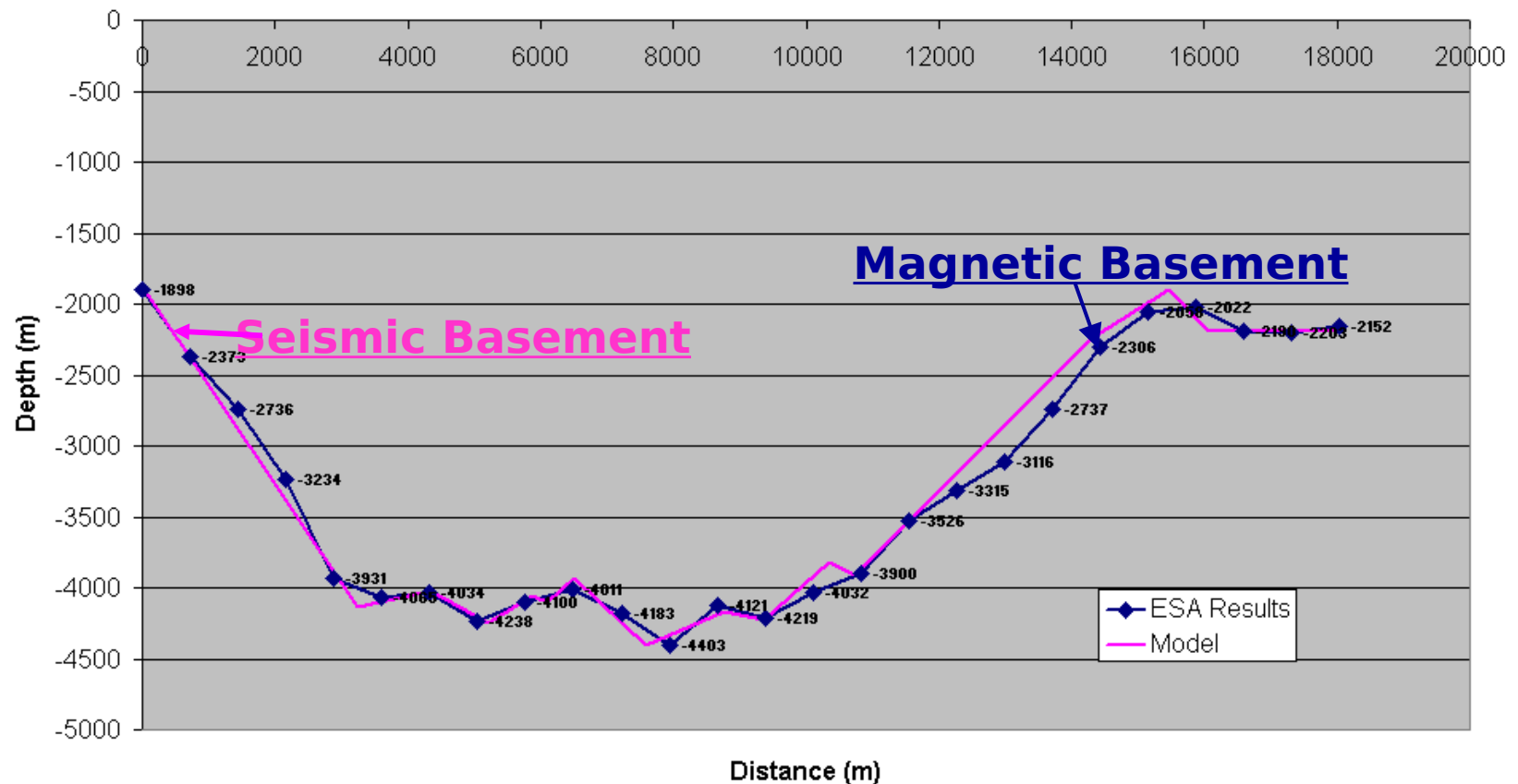


High Resolution Results

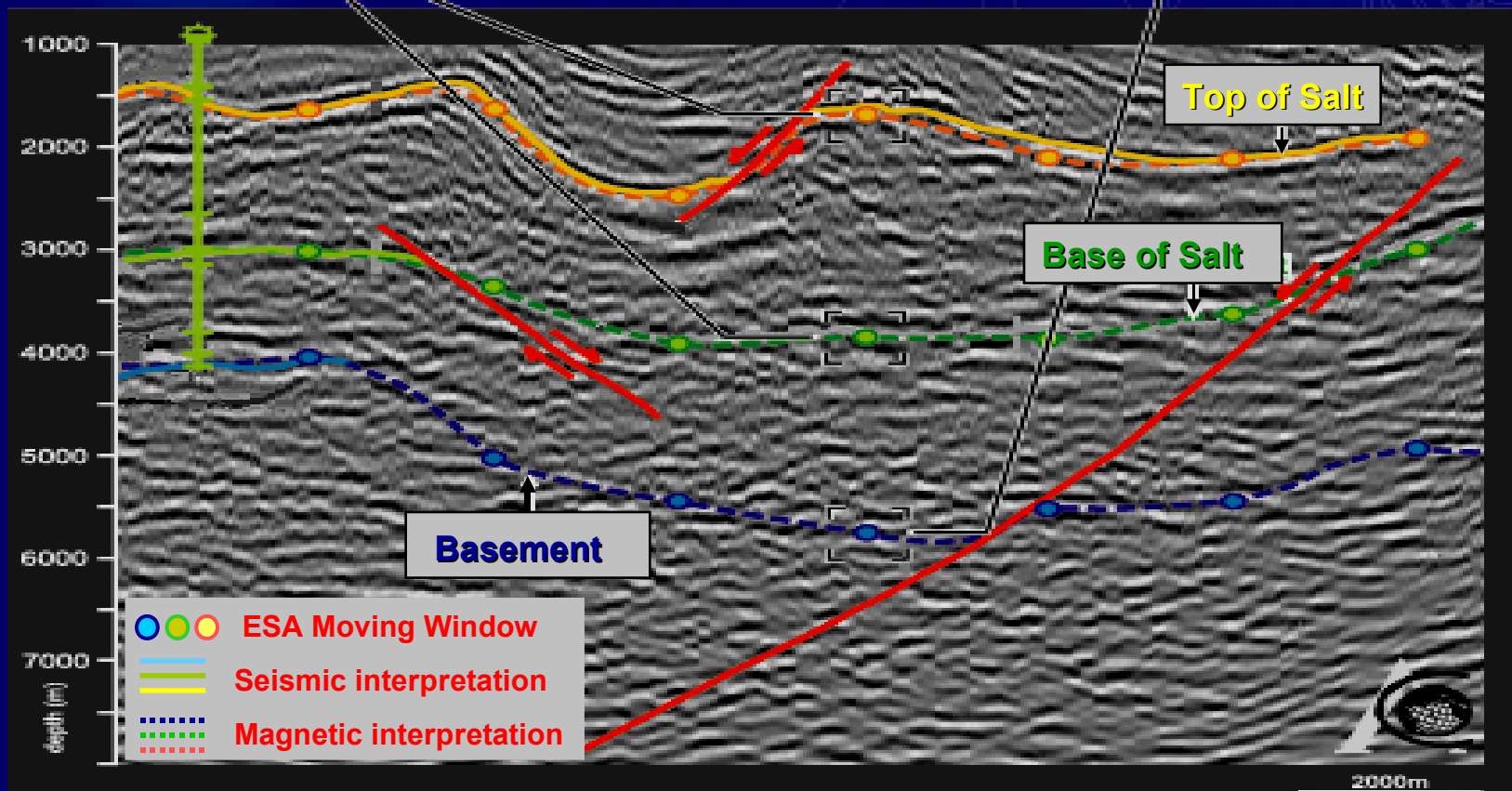
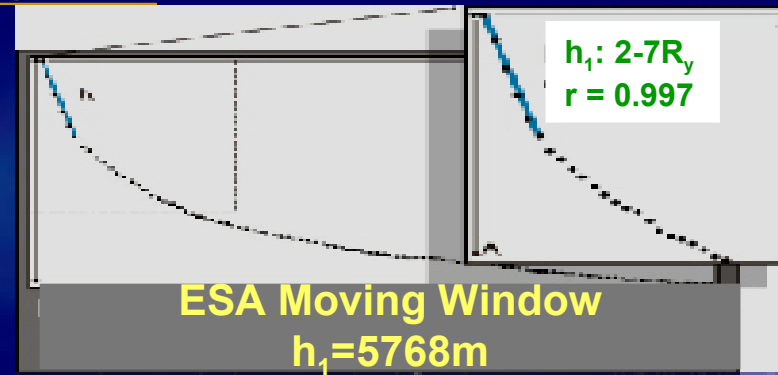
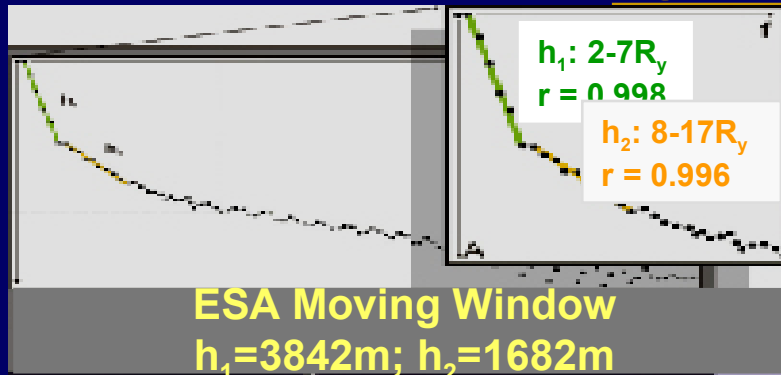


ESA 'Moving Window'

Depth to Magnetic Basement



Top Salt, Base Salt, Basement ESA-MW of TMI



ESA 'Moving Window' Technique

– Depth to Magnetic Horizons

- Semi-automatic Depth Calculation
- **AutoESA: Fully-automatic Depth Calculation**

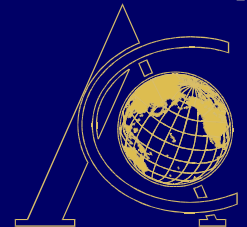
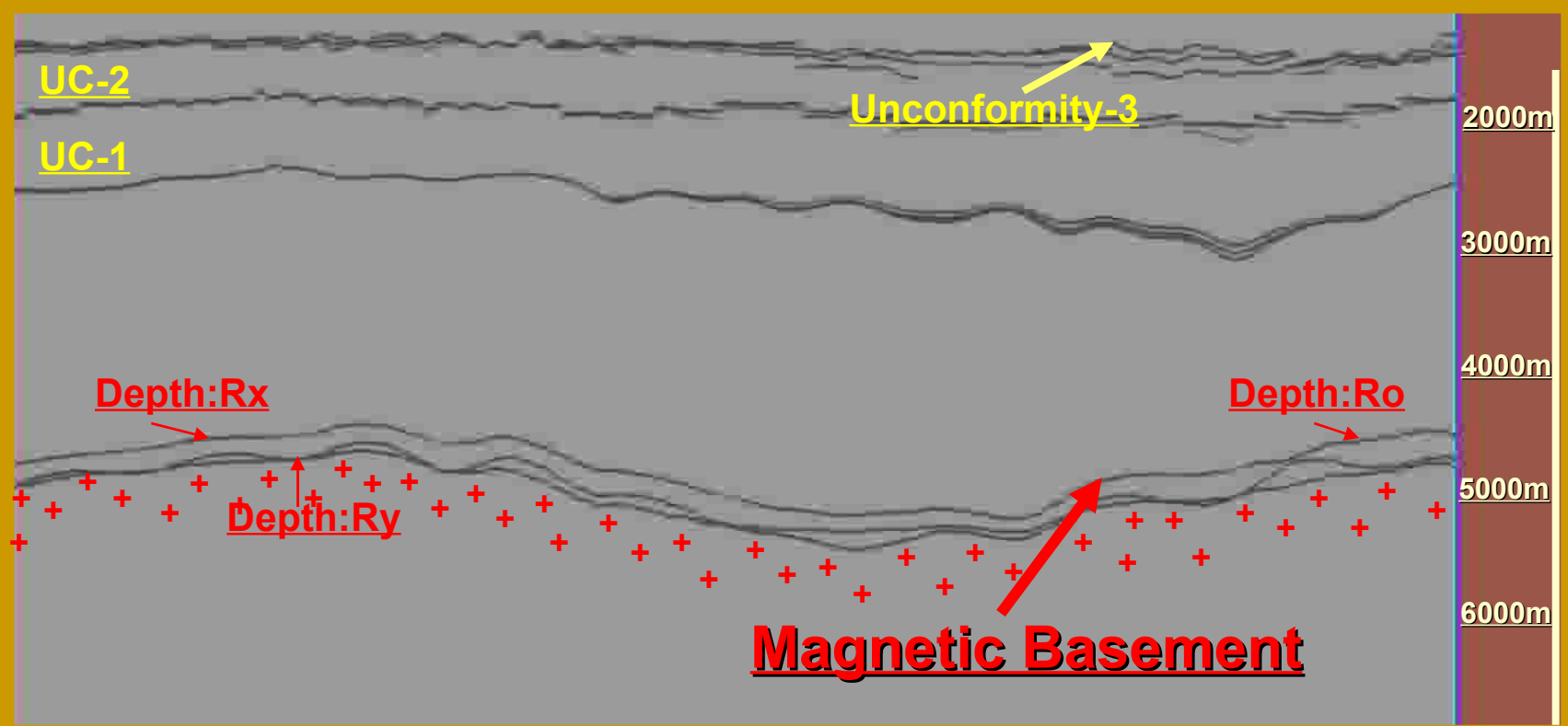


Basement & Intra-Sedimentary Horizons

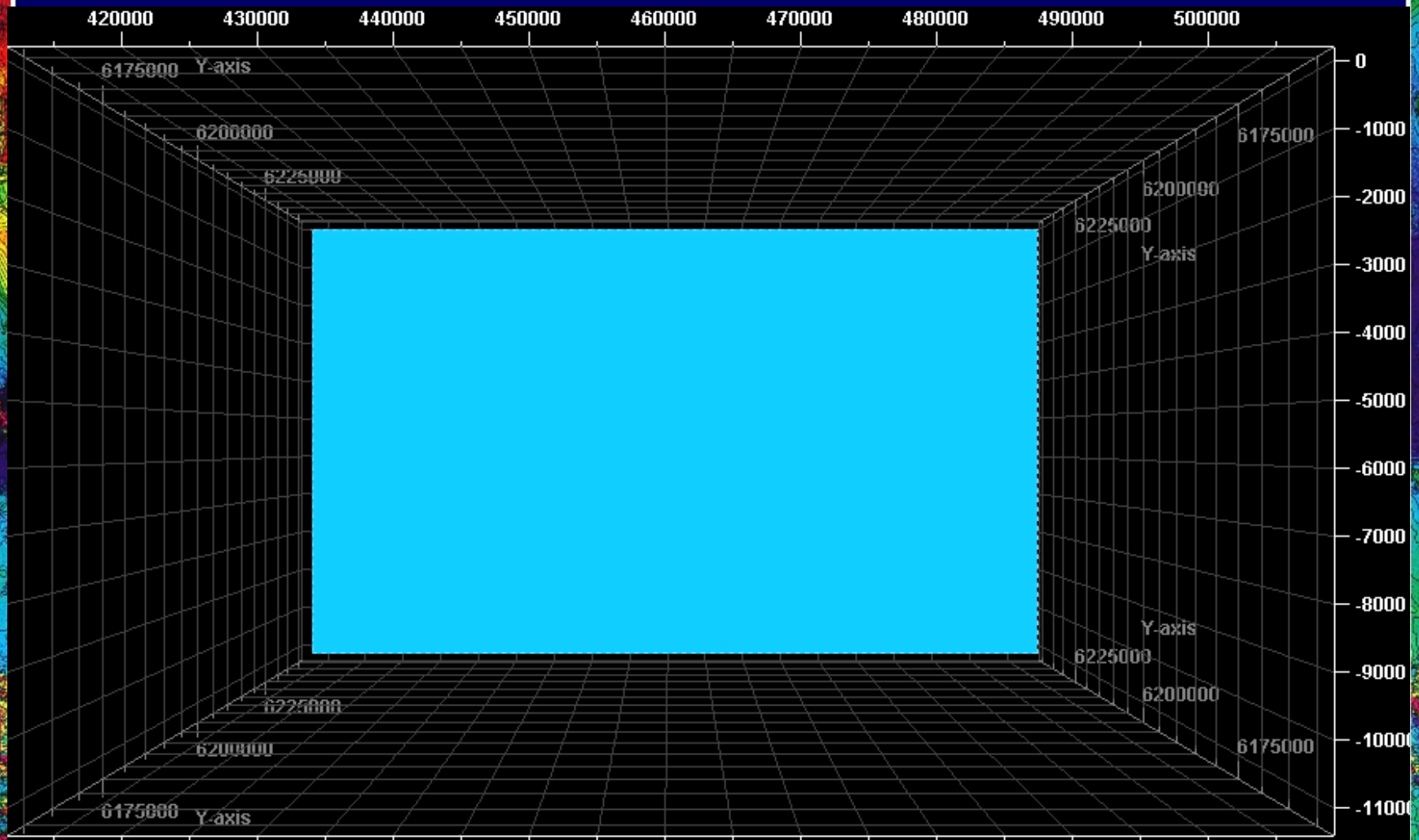
Detected by AutoESA of TMI

West

East

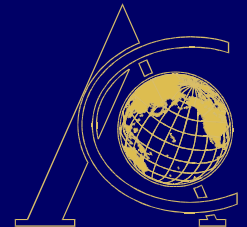


Comparison of AutoESA & Semi-Automatic Basement Interpretation EW Cross-Section in 3D-Volume



ESA 'Moving Window' Window Size

- **Calculated depth can depend on the window size**
 - too small window will not properly cover the anomalies
 - too large window will introduce averaging and possibly image a deeper structure
- **Optimal size of each window is detected using the Multi-Window Test**

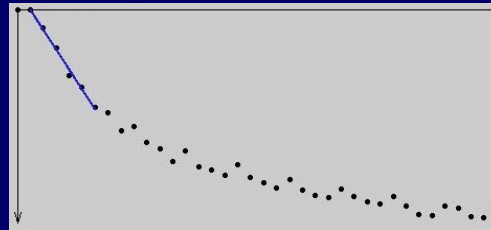
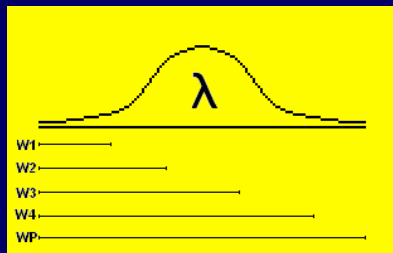
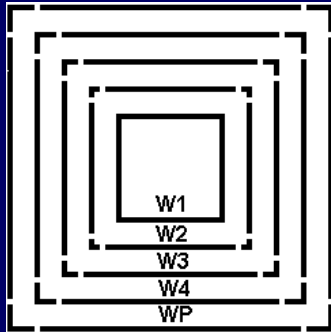


AutoESA Multi Window Test (AutoESA-MWT)

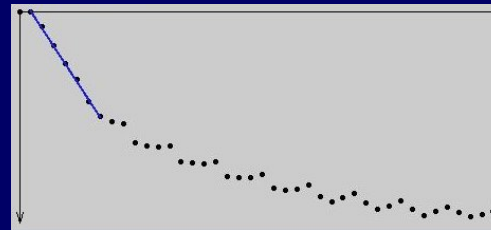


AutoESA MWT

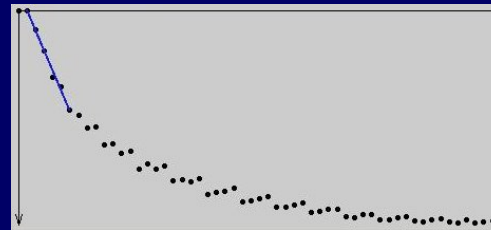
Depth Plateau Indicating Optimal Window Size



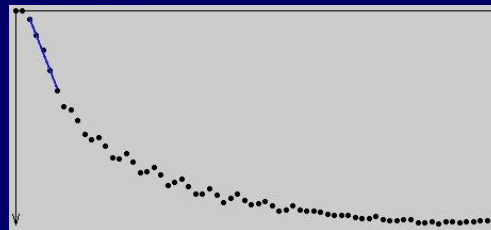
W1



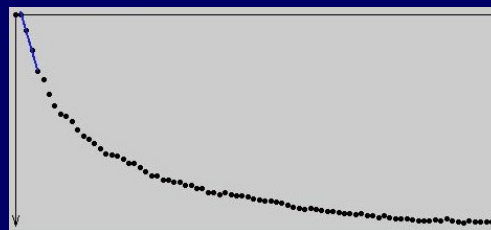
W2



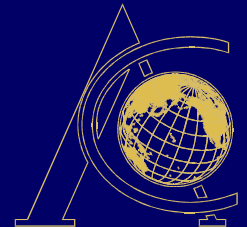
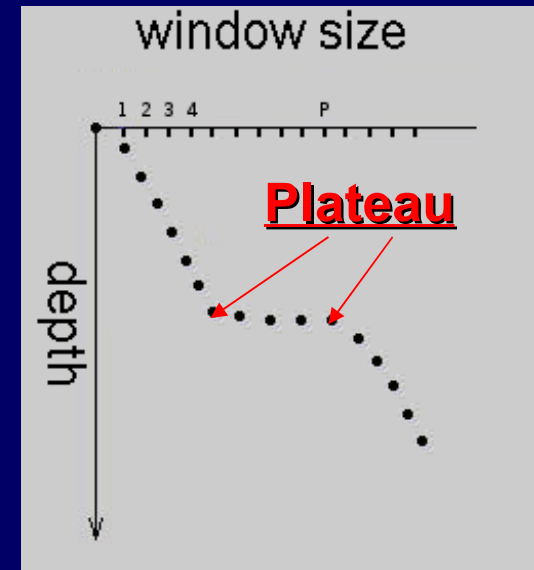
W3



W4



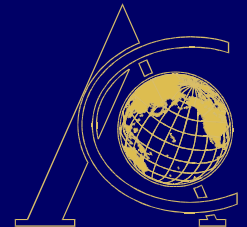
WP



AutoESA-MWT

Procedure

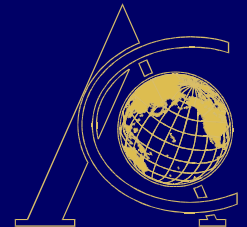
- **Spectra calculated using**
 - **Window Size**
 - Minimum, for example 20gc
 - Maximum related to targeted depth
 - Increment usually 2gc
 - **MWT stations**
 - Regular grid e.g. 1x1gc
 - Profiles across the study area



AutoESA-MWT

Procedure

- For each MWT station is computed
 - Window density of plateaus
 - Spectra shape parameters at each window size
 - Statistical goodness of fit



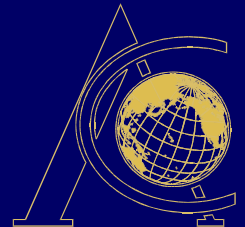
AutoESA-MWT

Window Density

- For each depth plateau

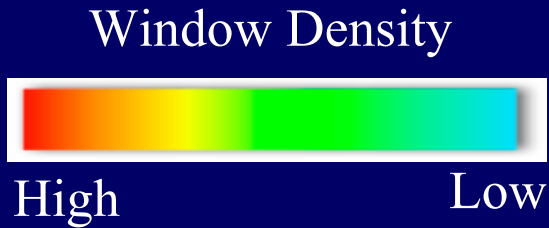
$$\text{Window Density} = \text{Number of windows forming plateau}$$

- High Window Density implies robust depth estimation

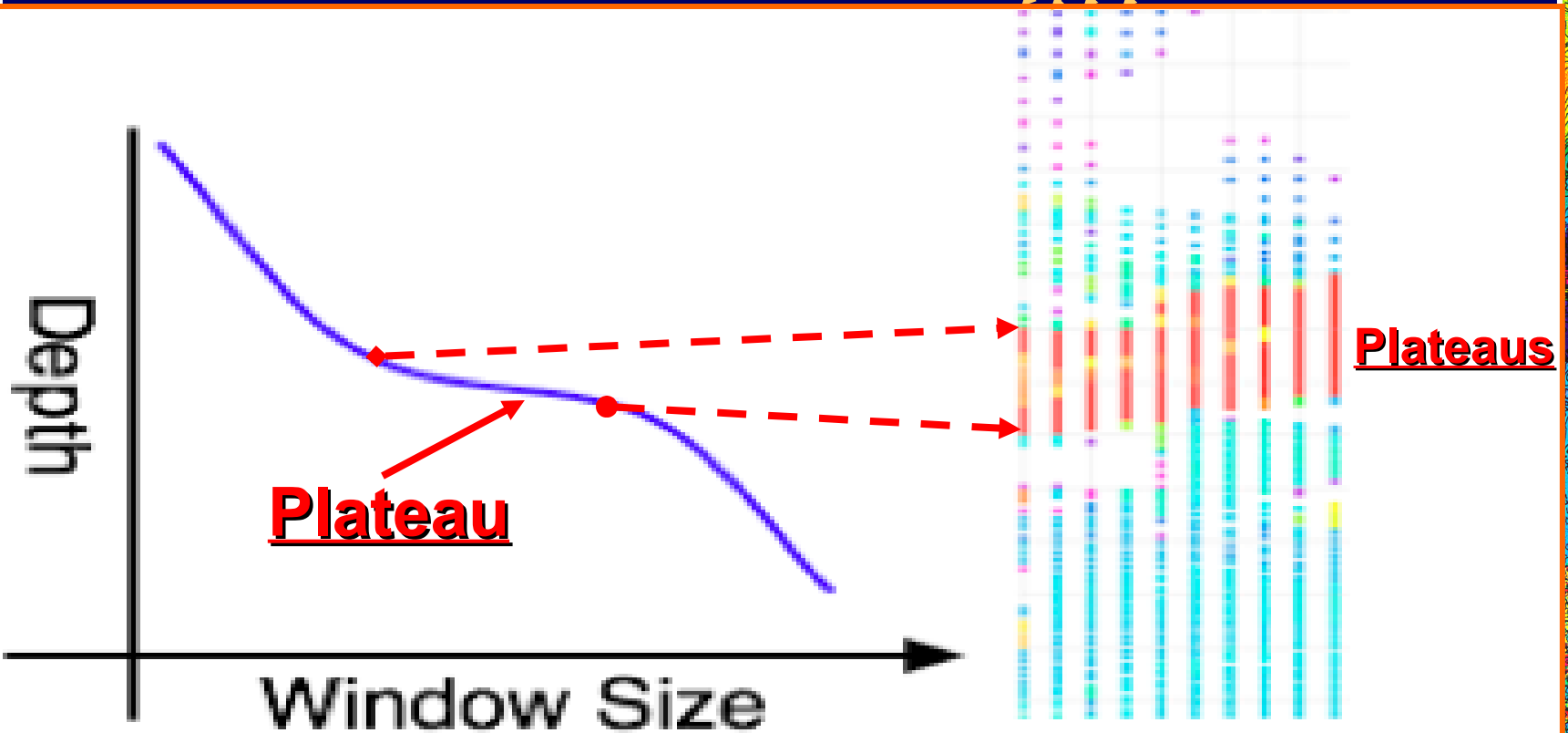


AutoESA MWT

Window Density of Plateaus



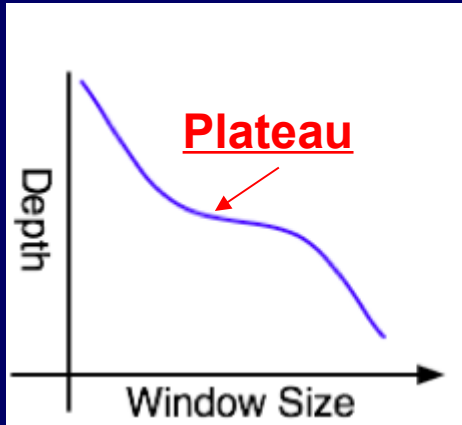
MWT Stations



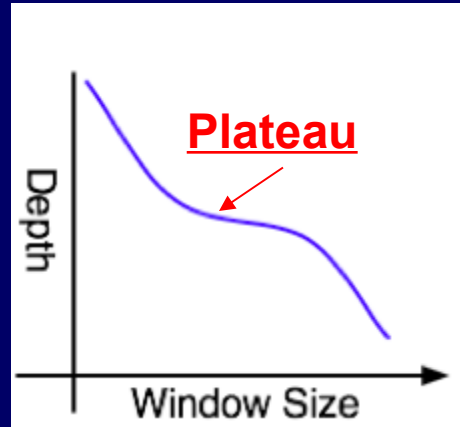
AutoESA MWT

Window Density of Plateaus along Profile Images Horizon

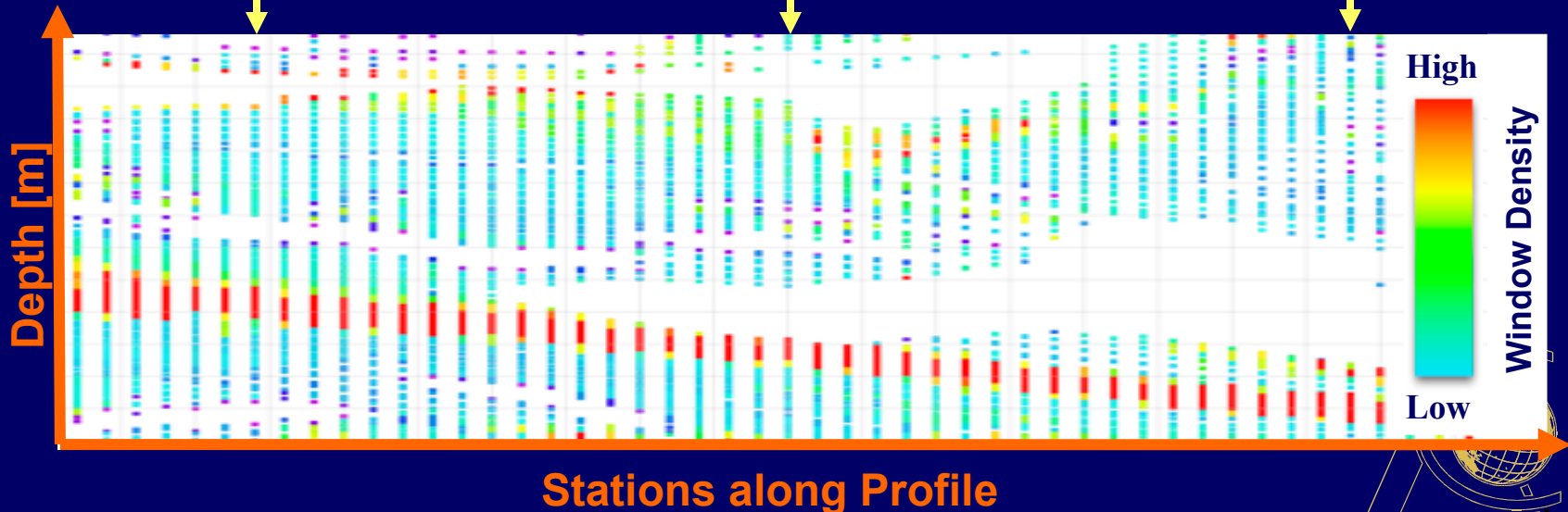
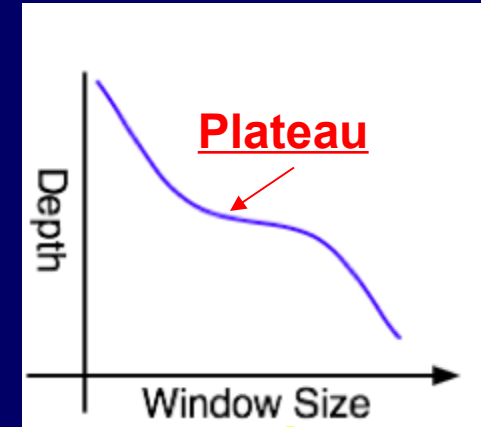
Station-7



Station-25



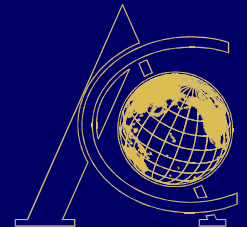
Station-44



AutoESA-MWT

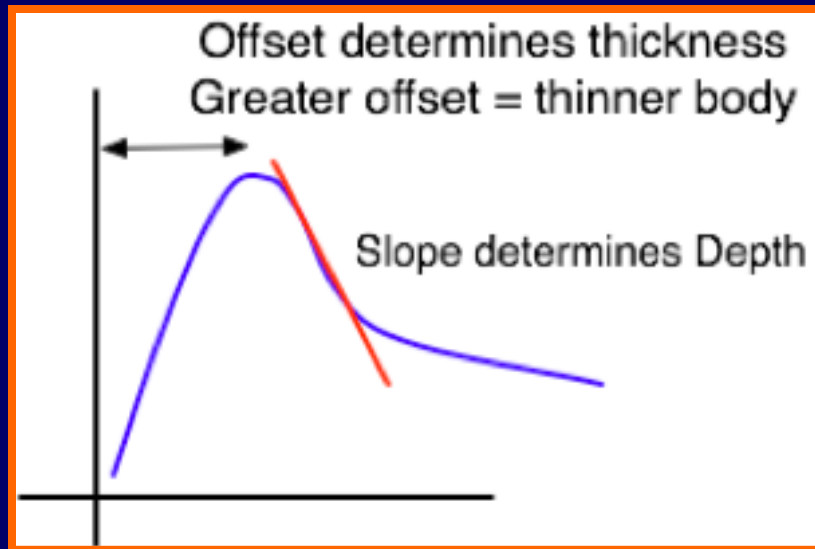
Procedure

- For each MWT station is computed
 - Window density of plateaus
 - Spectra shape parameters at each window size
 - Statistical goodness of fit

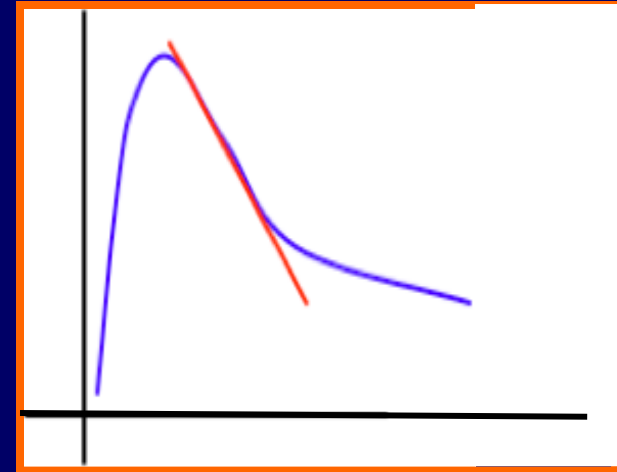


AutoESA MWT

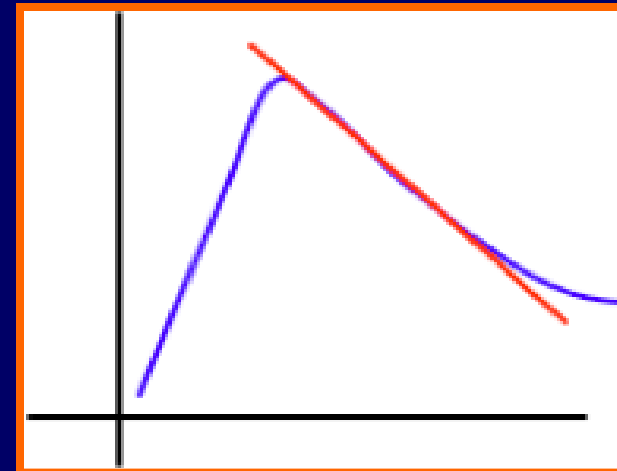
Analysis of Spectra Shape



**Geometry of Body
influences Spectrum Shape**



Thick deep body

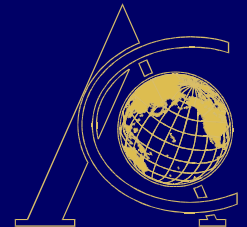


Shallow thin body

AutoESA-MWT

Analysis of Spectra Shape

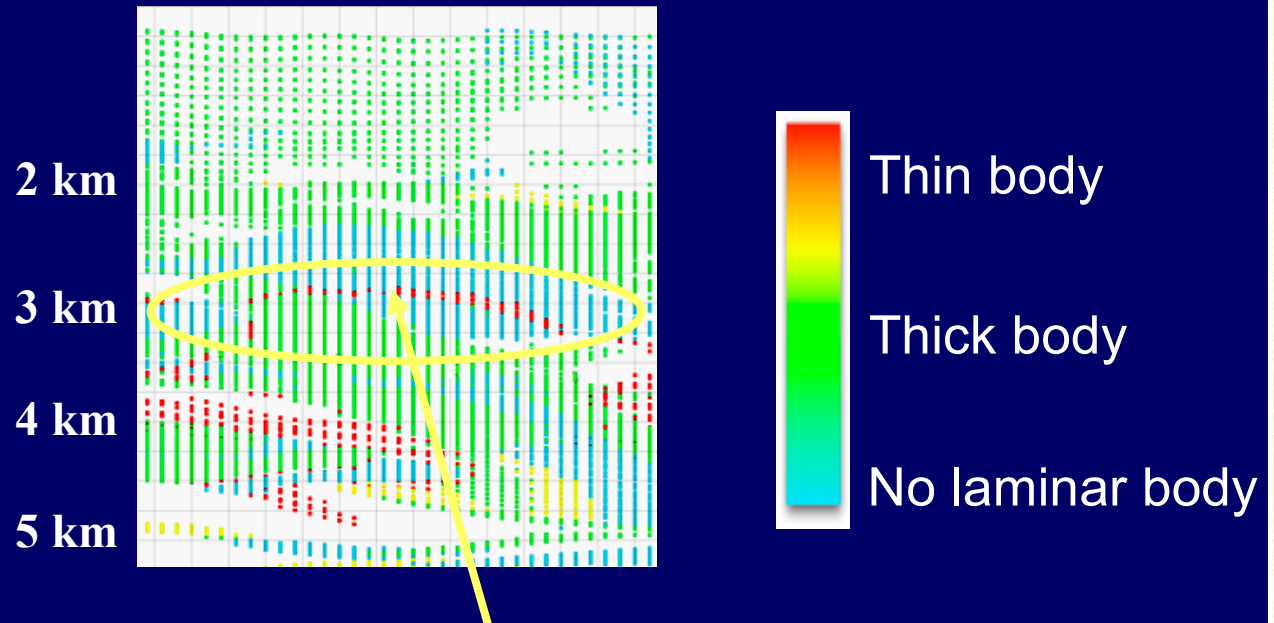
- For each spectrum:
 - the shape is quantified
 - depth of the 1st slope is computed
- Spectra shape parameter is plotted vs computed depth
 - For visualization along a profile, high values of the spectra shape parameter are represented by hot colours
- High values of the spectra shape parameter correspond to thin bodies



AutoESA-MWT

Analysis of Spectra Shape

**Spectra Shape Parameter vs Depth
along Profile**

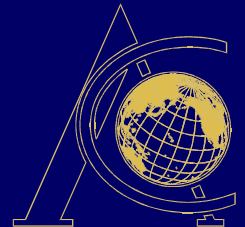


Detecting Laminar Bodies



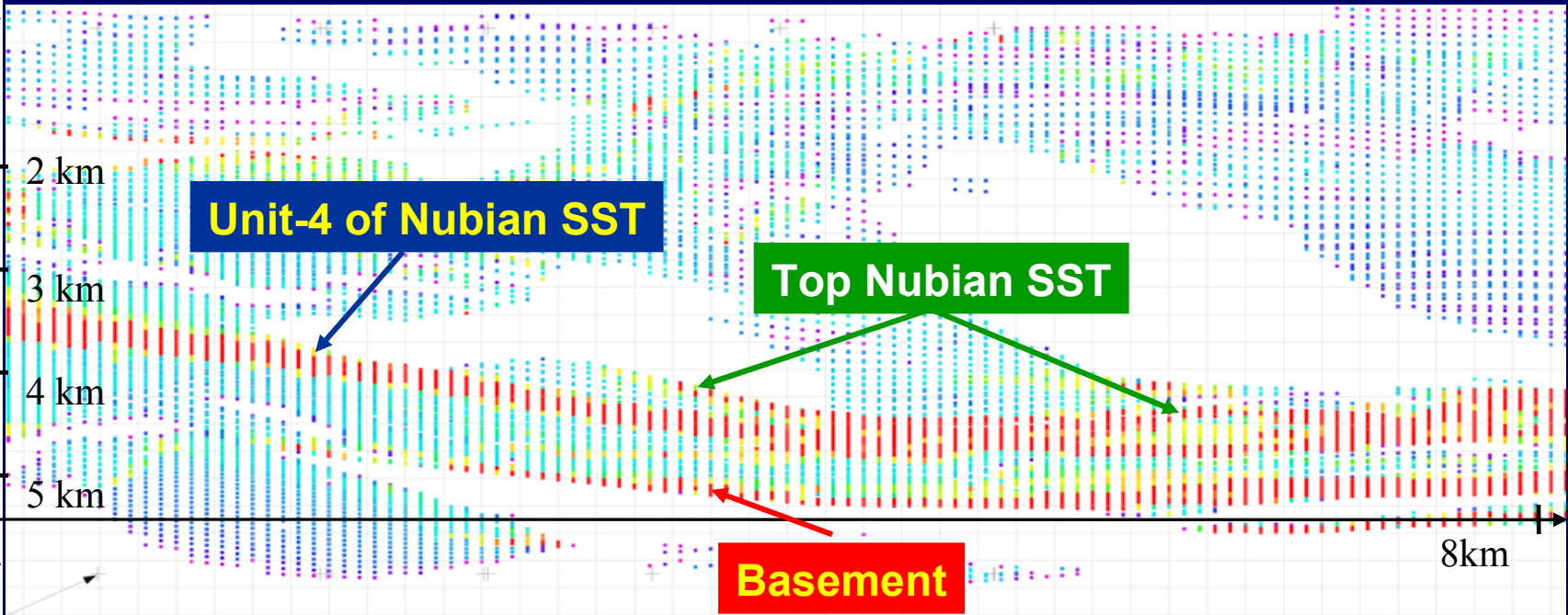
Outline

- Introduction
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 - Basement
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 - Unit 4 of Nubian SST
- Comparison with Well and Seismic data
- Conclusions



AutoESA-MWT: Horizon Mapping Window Density

Sirt Basin: Libya



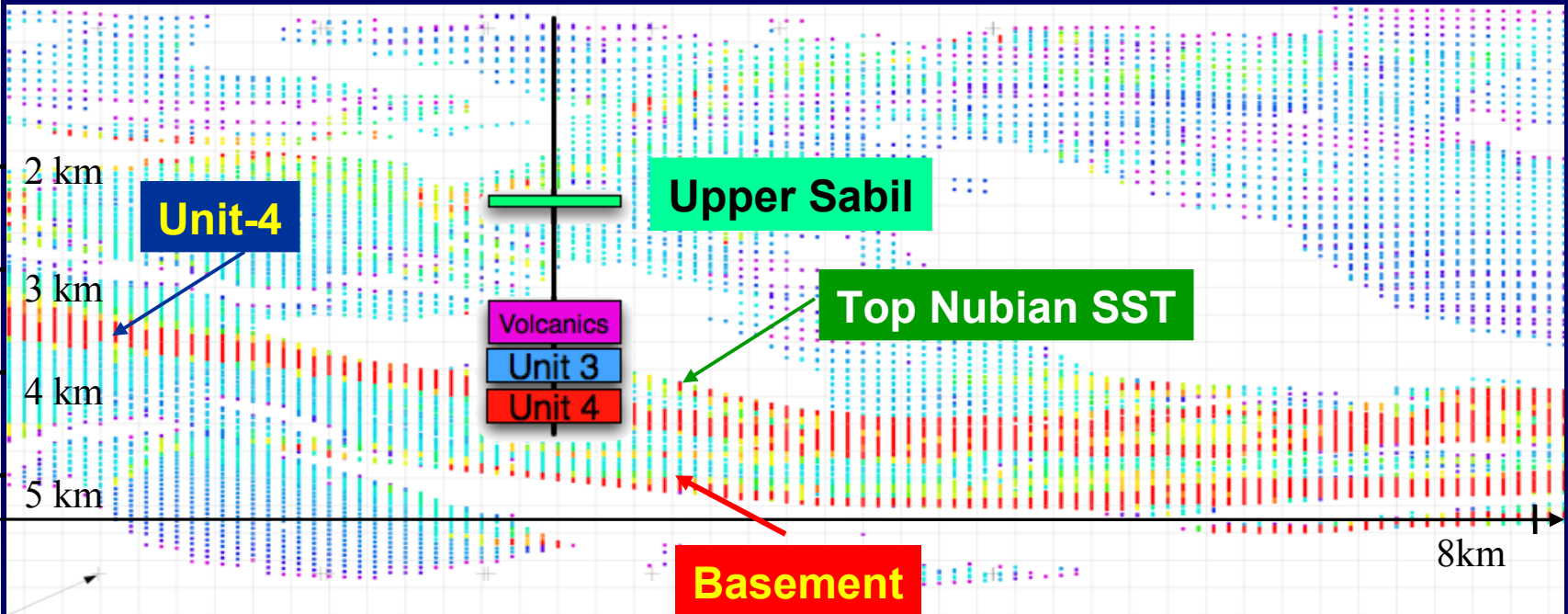
Window Density



AutoESA-MWT: Horizon Mapping

Window Density

Sirt Basin: Libya



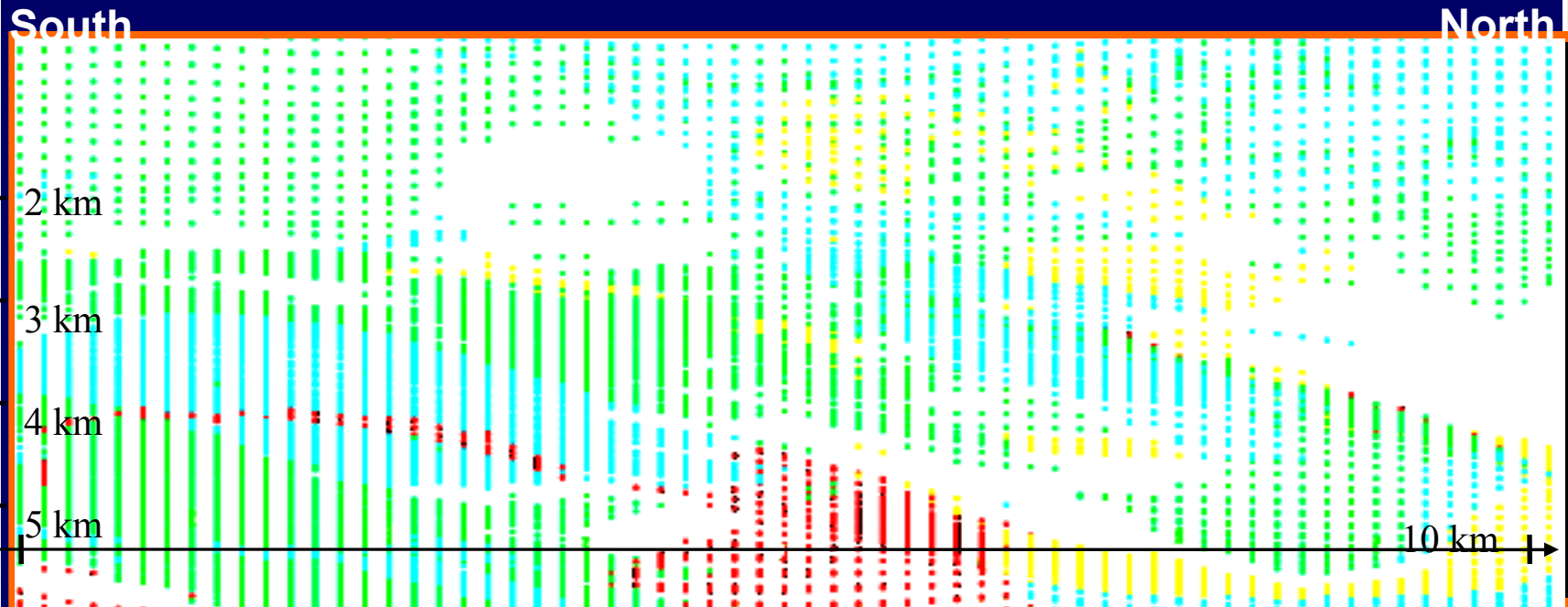
Window Density



AutoESA-MWT: Horizon Mapping

Spectra Shape: Detecting Laminar Bodies

Sirt Basin: Libya



Thin body

Thick body

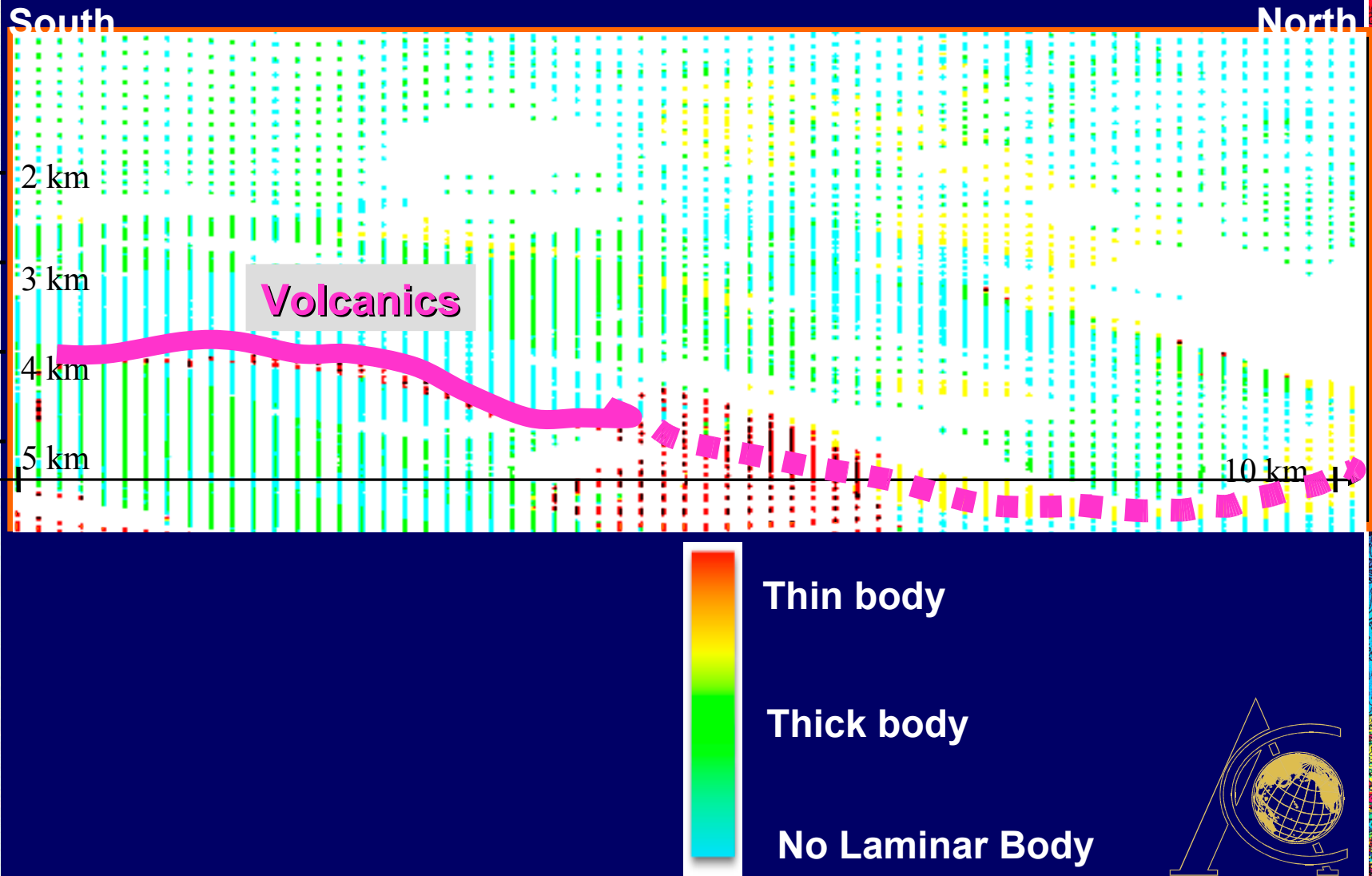
No Laminar Body



AutoESA-MWT: Horizon Mapping

Spectra Shape: Detecting Laminar Bodies

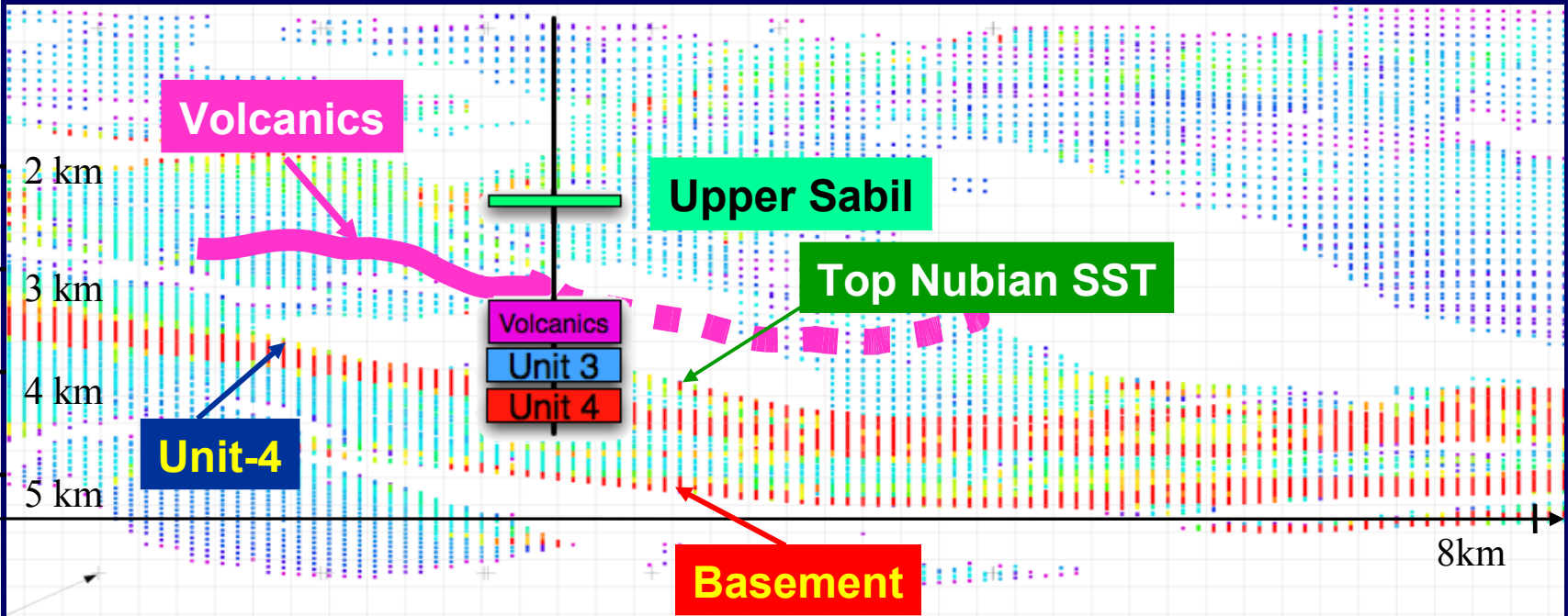
Sirt Basin: Libya



AutoESA-MWT: Horizon Mapping

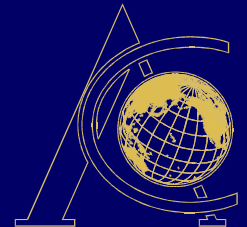
Window Density

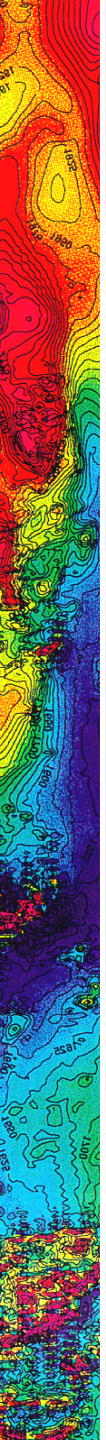
Sirt Basin: Libya



High window density

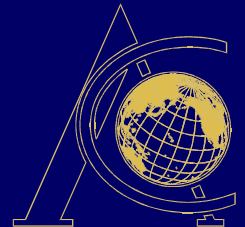
Low window density





ESA-MW

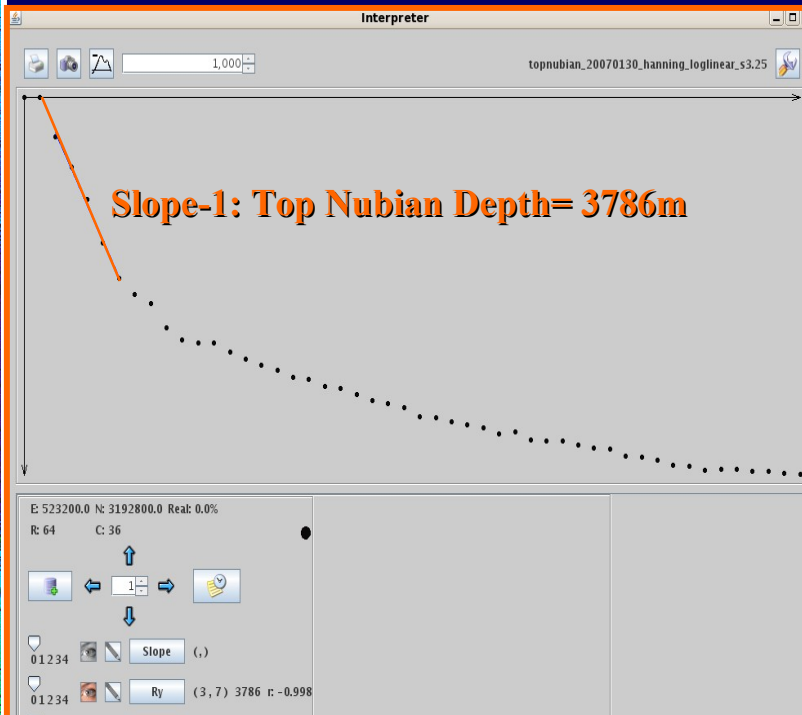
Top Nubian Horizon



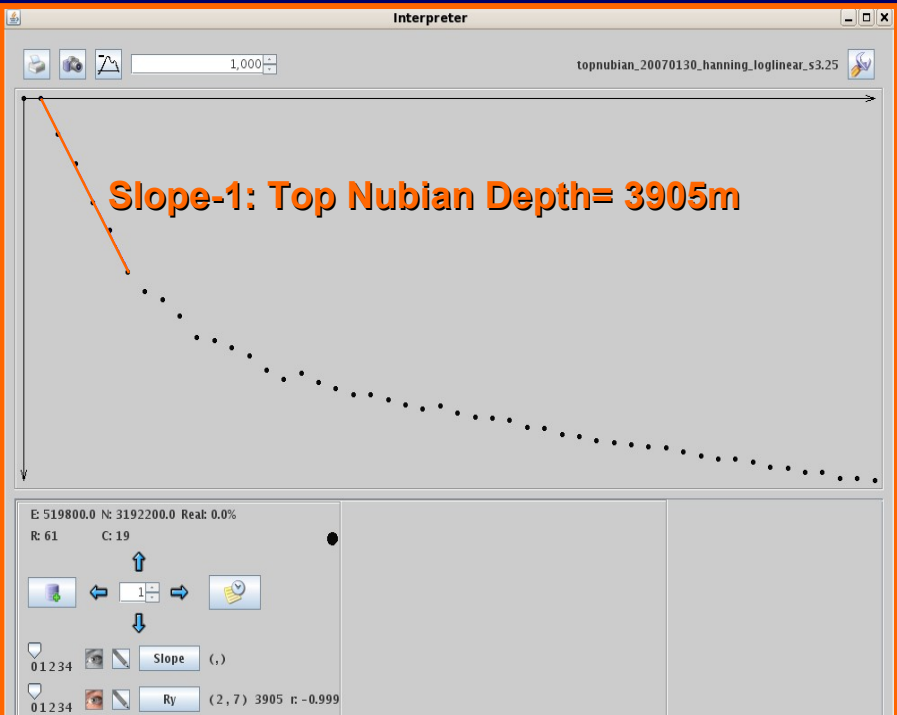
Top Nubian: ESA-MW

Example of Interpreted Spectra Semi-automatic Depth Calculation

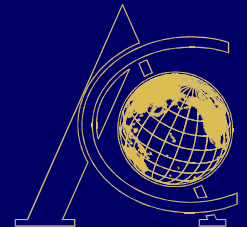
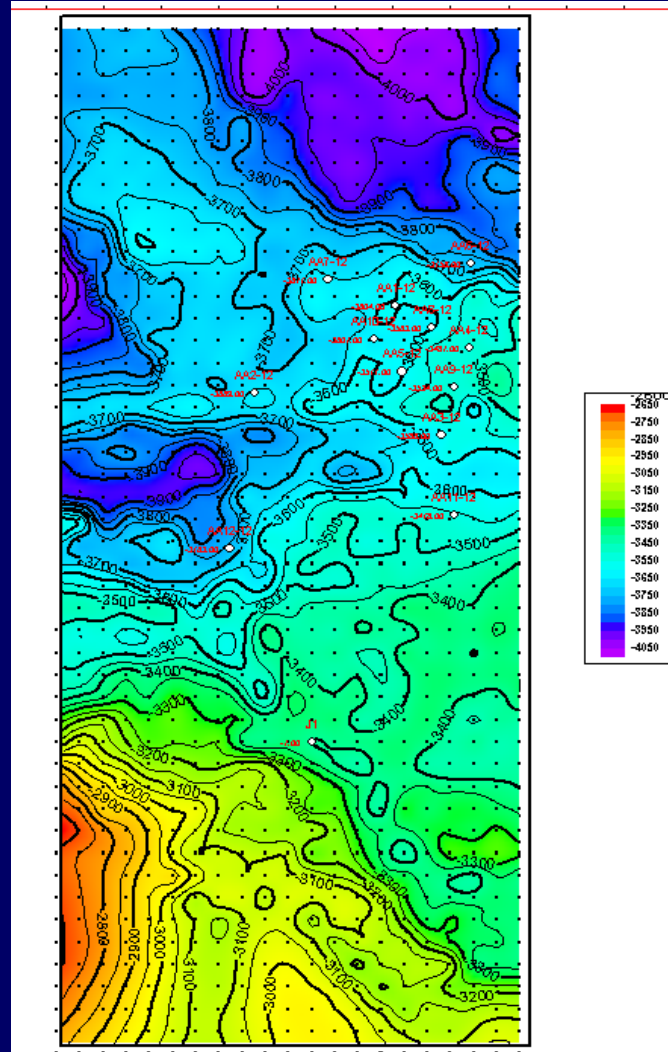
Location-1



Location-2

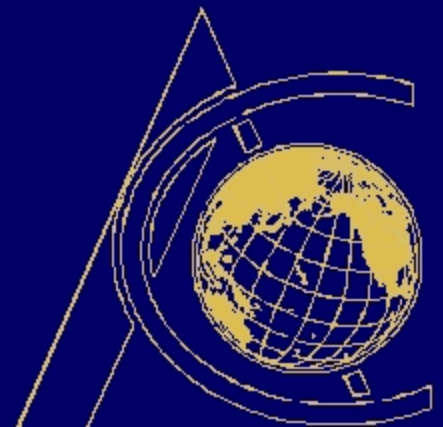


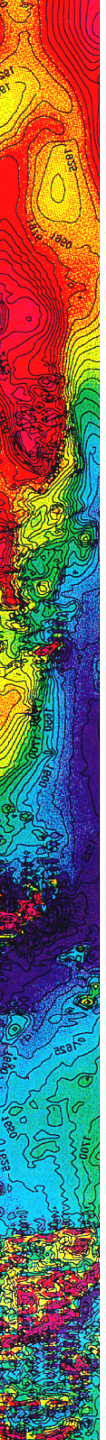
Top Nubian: ESA-MW



Top Nubian Horizon

ESA-MW 3D Visualisation





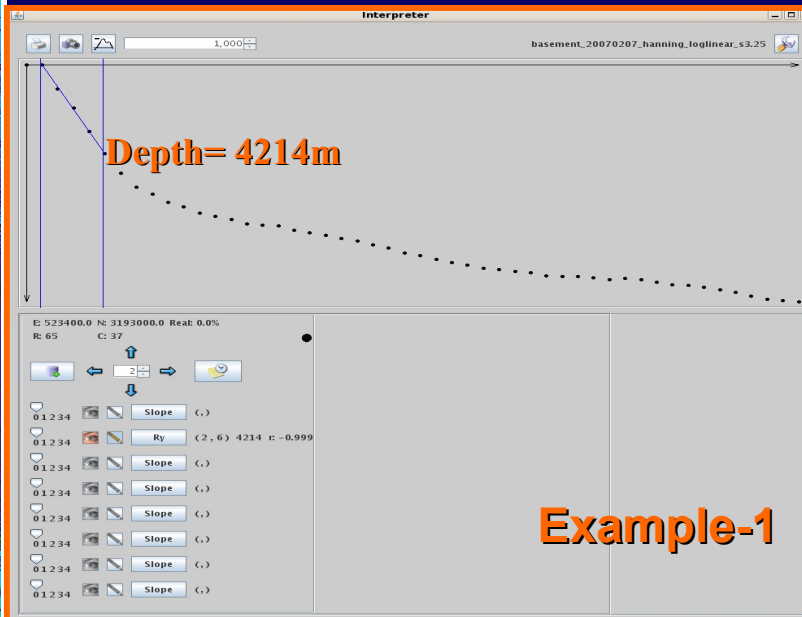
ESA-MW

Basement Horizon

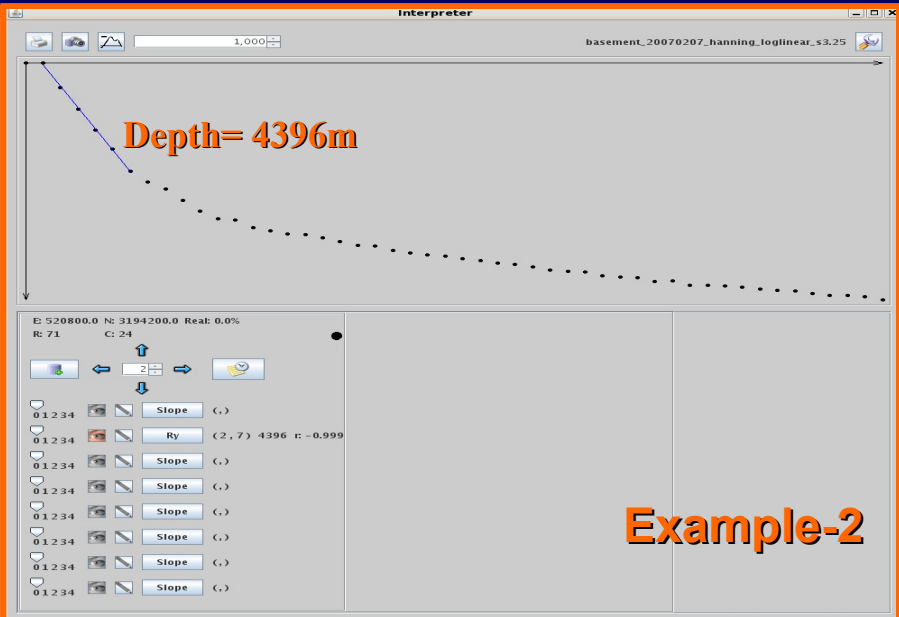


Magnetic Basement: ESA-MW

Example of Interpreted Spectra Semi-automatic Depth Calculation



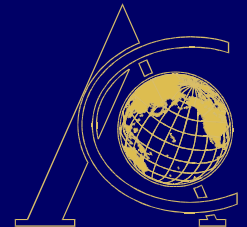
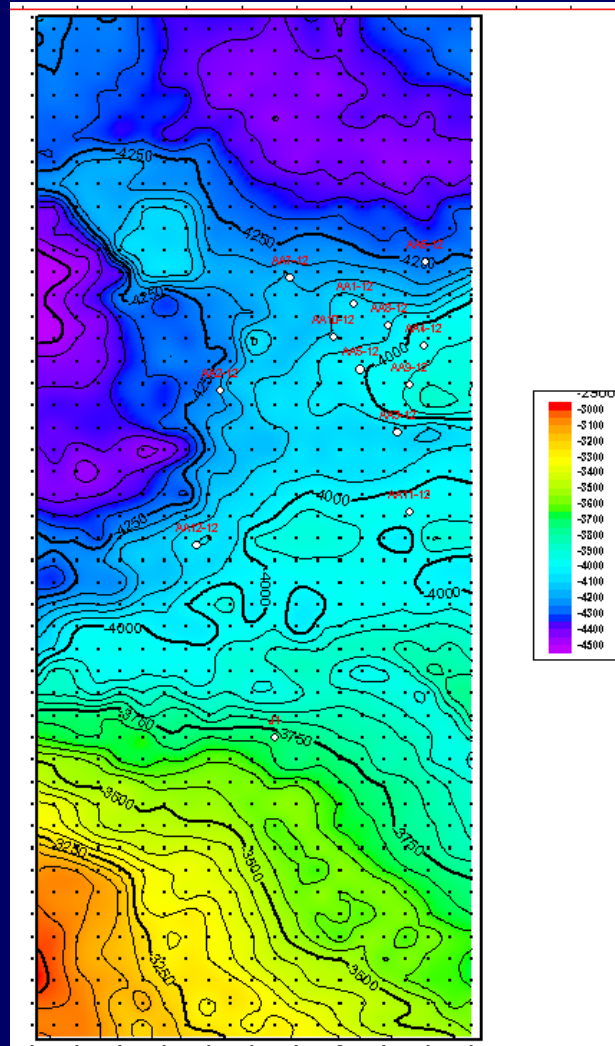
Example-1

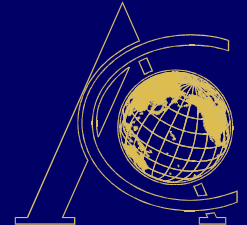
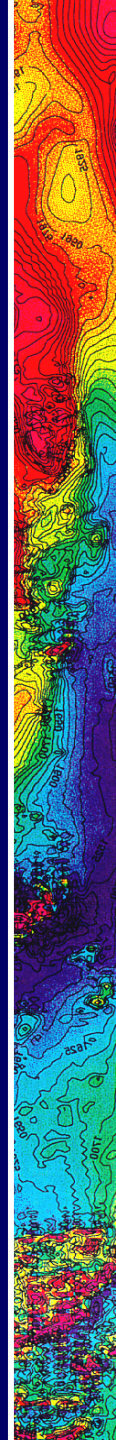


Example-2

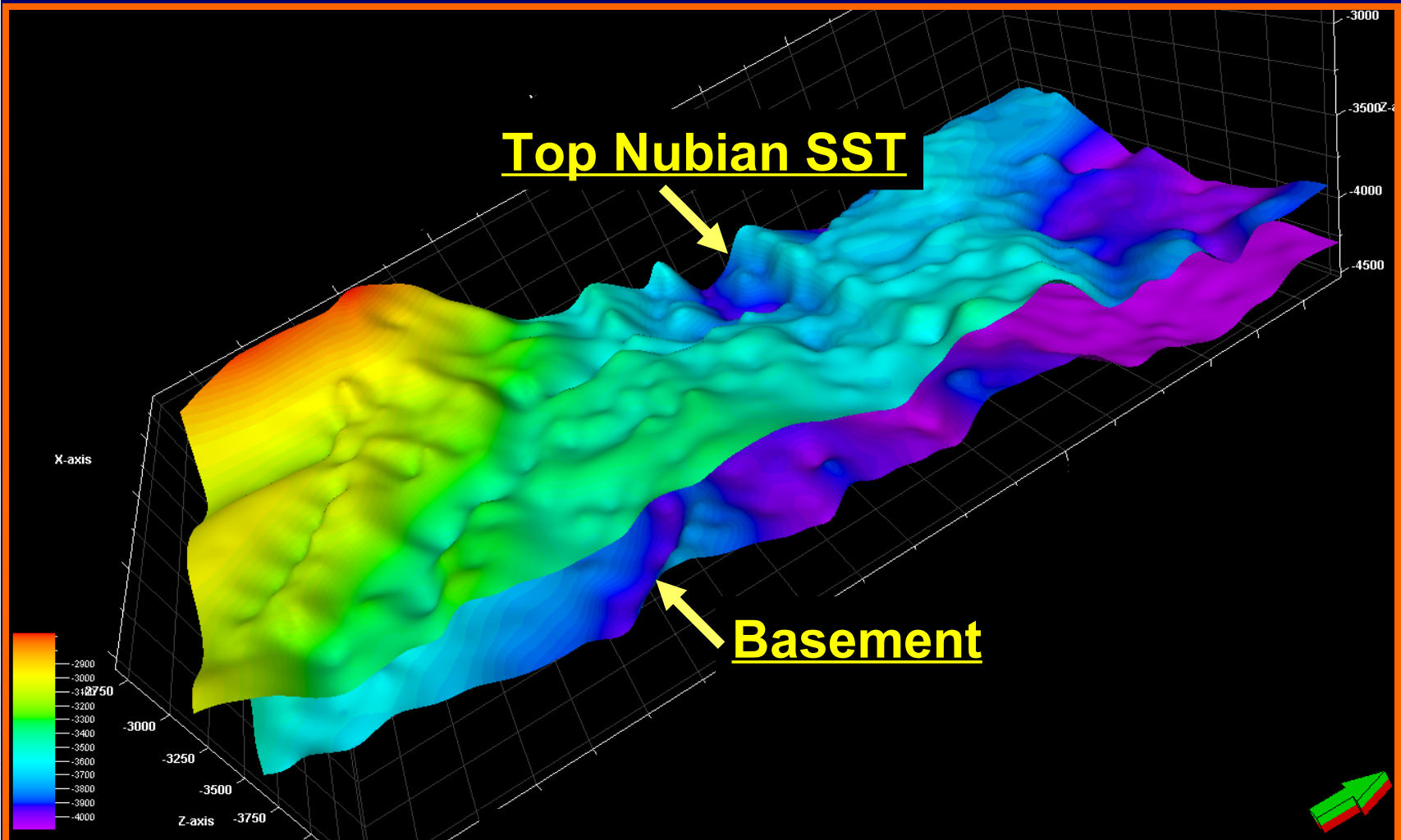


Basement: ESA-MW



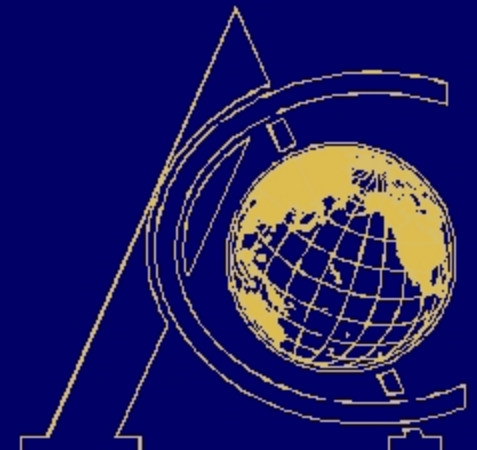


Horizon 3D View



Basement & Top Nubian SST Depth Calculation from ESA-MW

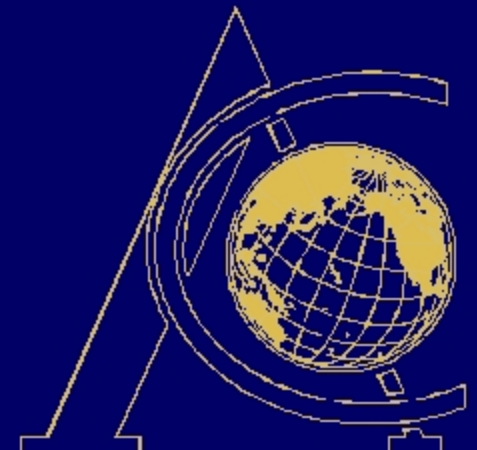
Semi-automatic: Manual Picks - Points
3D Visualisation



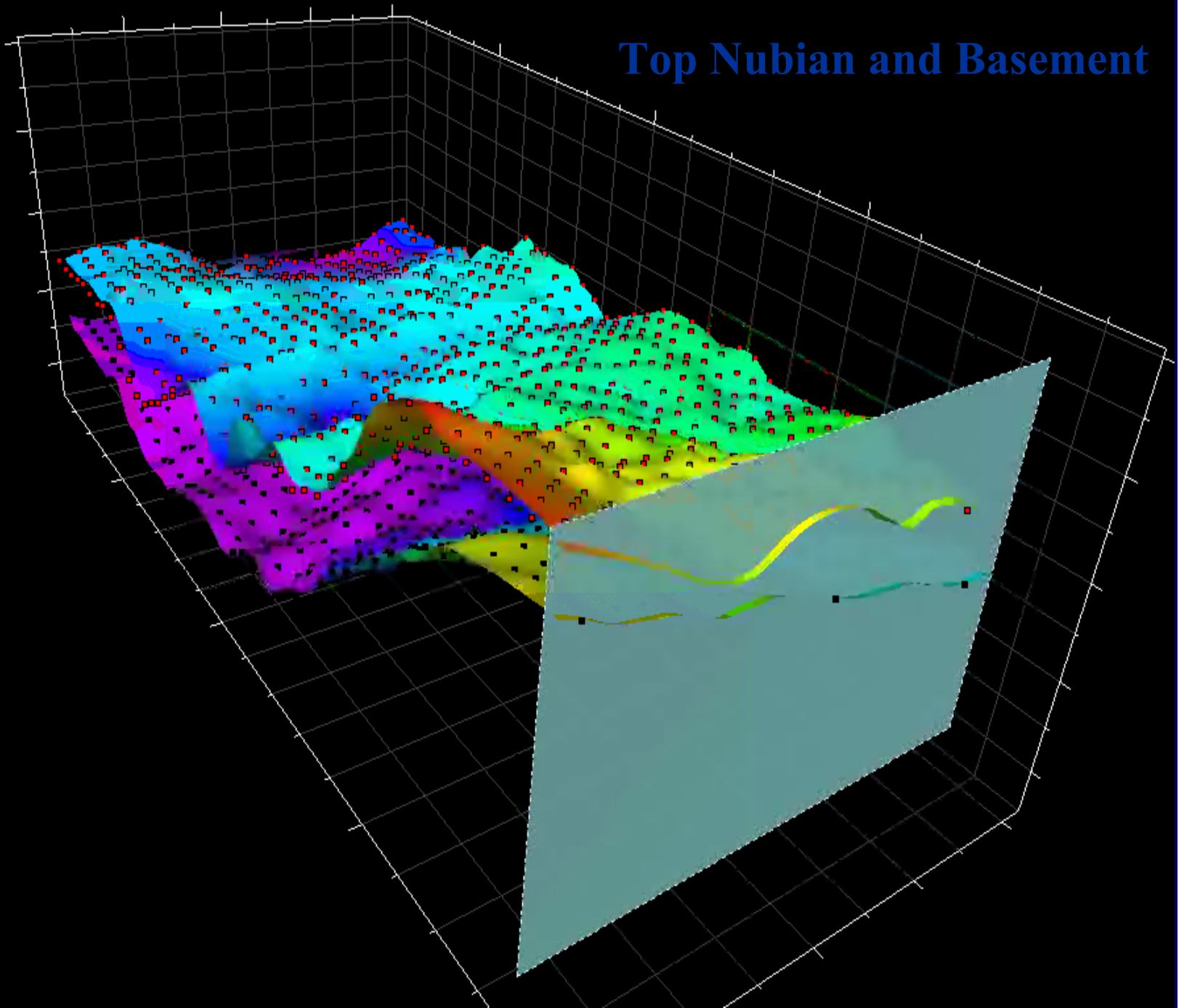
Top Nubian & Basement

ESA-MW

3D Visualisation

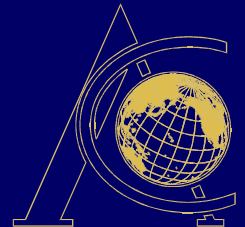


Top Nubian and Basement



Outline

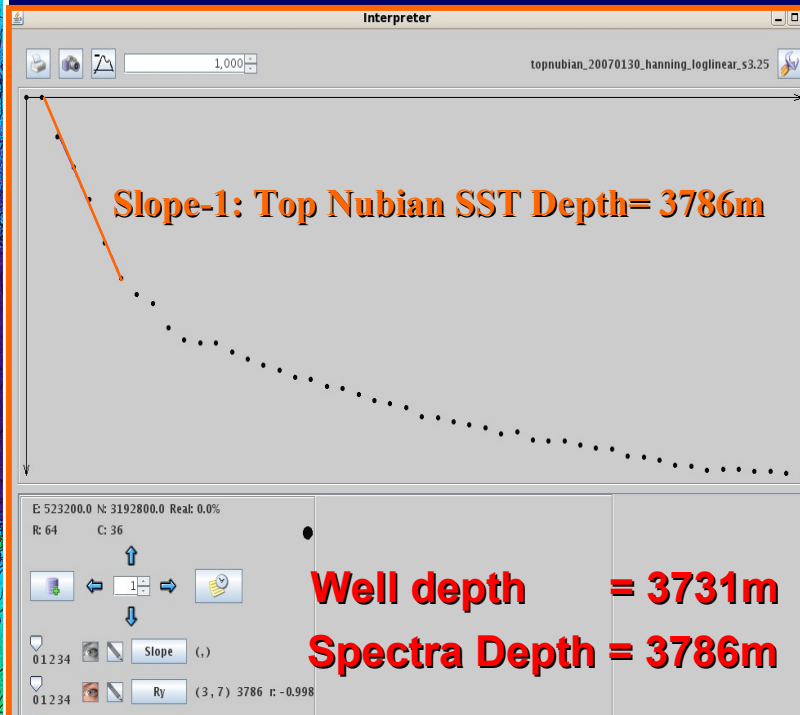
- Introduction
- Methodology
- Horizon Interpretation
 - Basement
 - Top Nubian
 - Unit 4 of Nubian SST
- Comparison with Well and Seismic data
- Conclusions



Well Correlation: ESA-MW

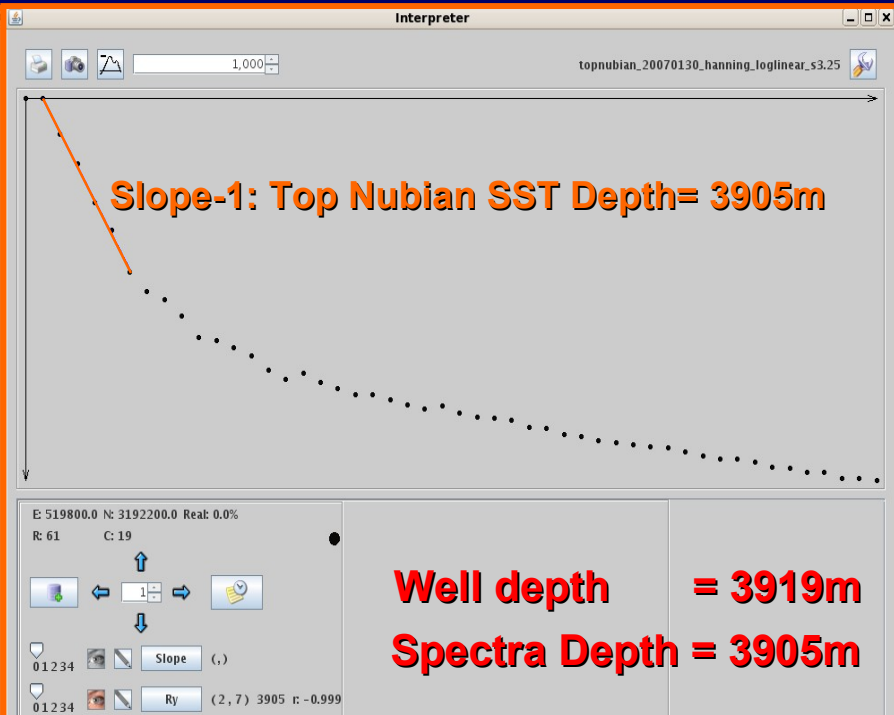
Example of Spectra & Well Depth Semi-automatic Depth Calculation

Well-1

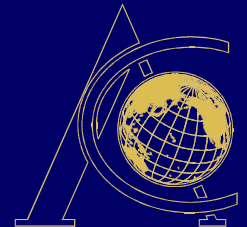


Well depth = 3731m
Spectra Depth = 3786m

Well-2

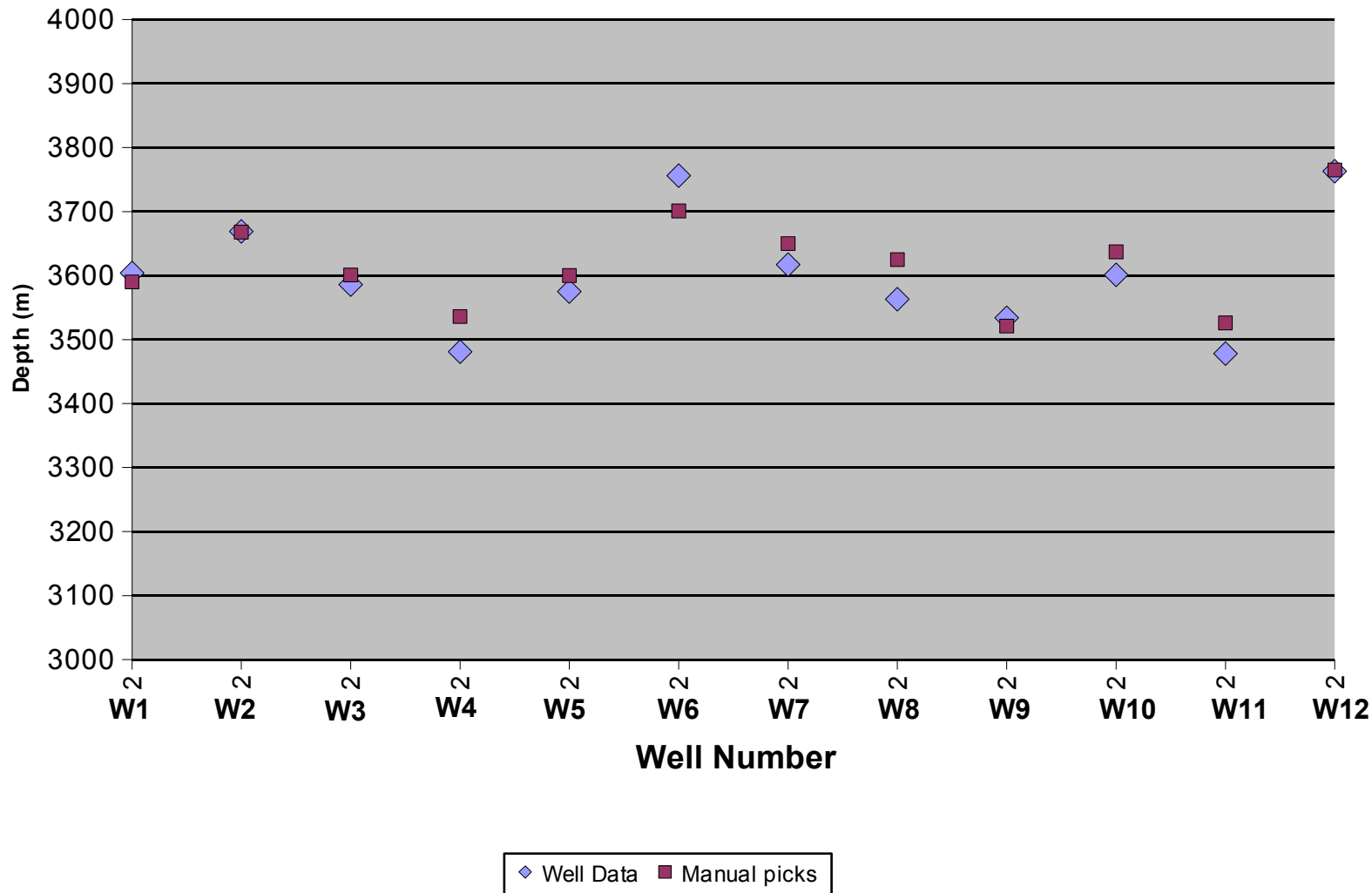


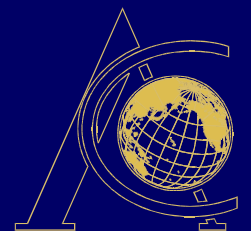
Well depth = 3919m
Spectra Depth = 3905m



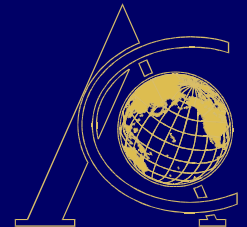
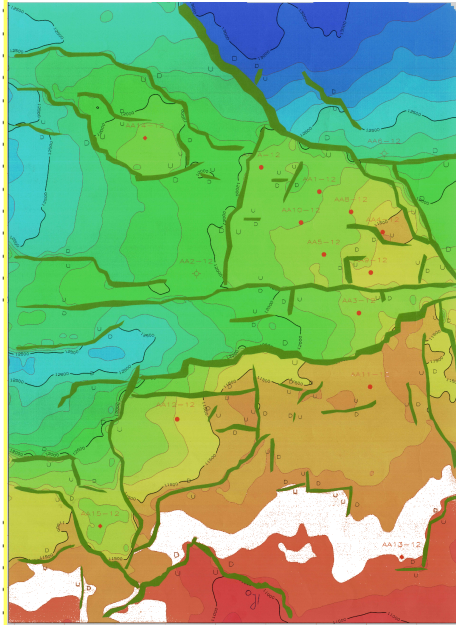
Well Depth vs Spectra Depth

Example from Sirt Basin: Top Nubian



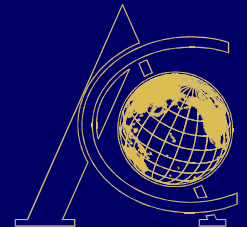
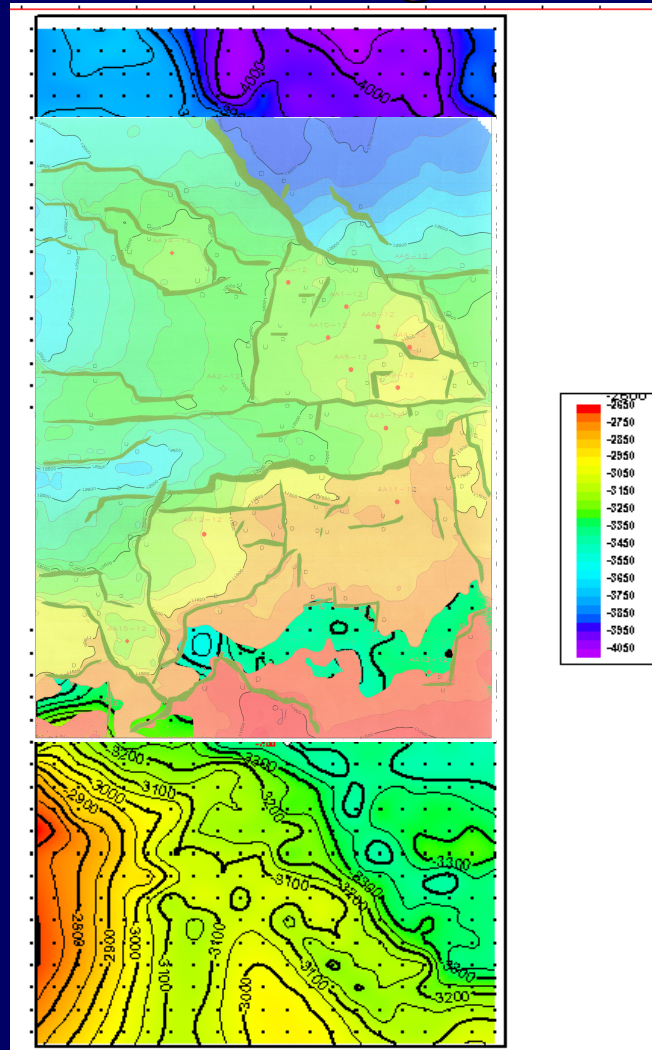


Top Nubian: Seismic



Top Nubian

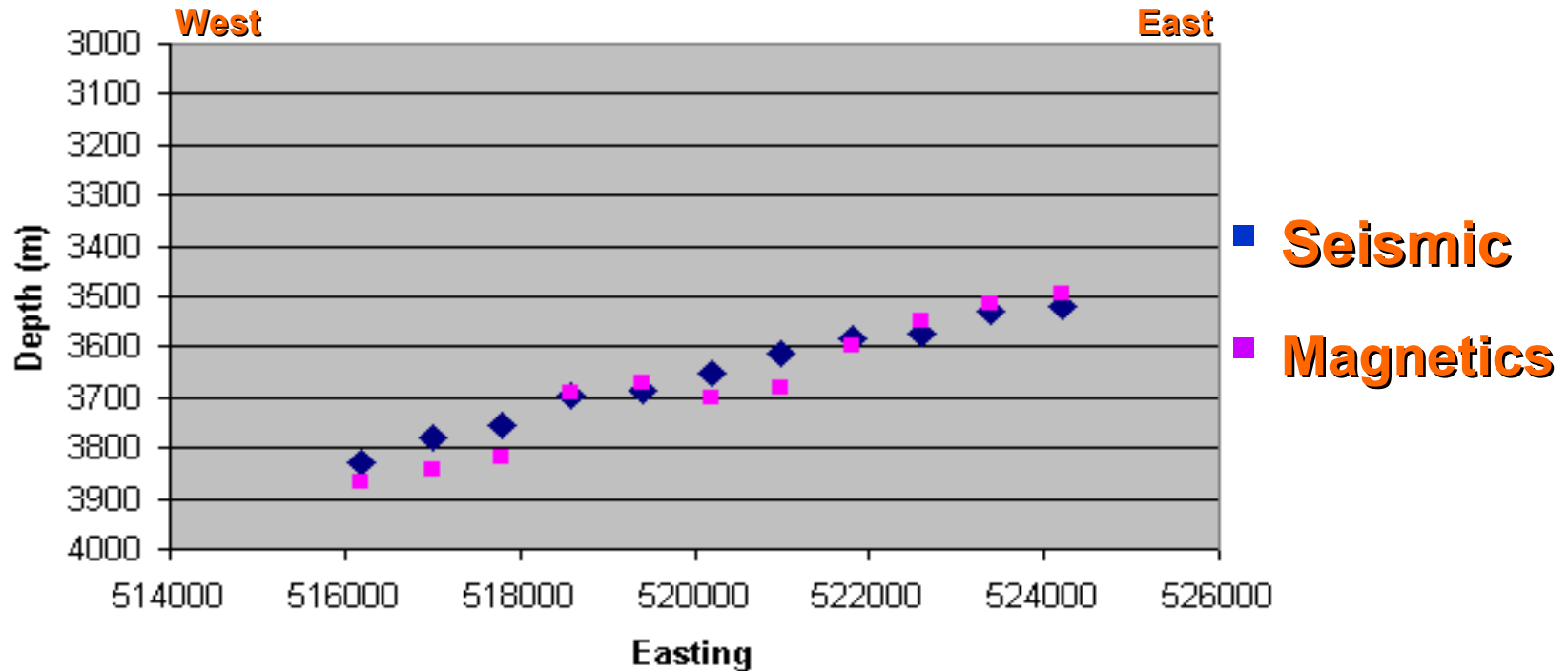
Seismic vs Magnetics



Top Nubian

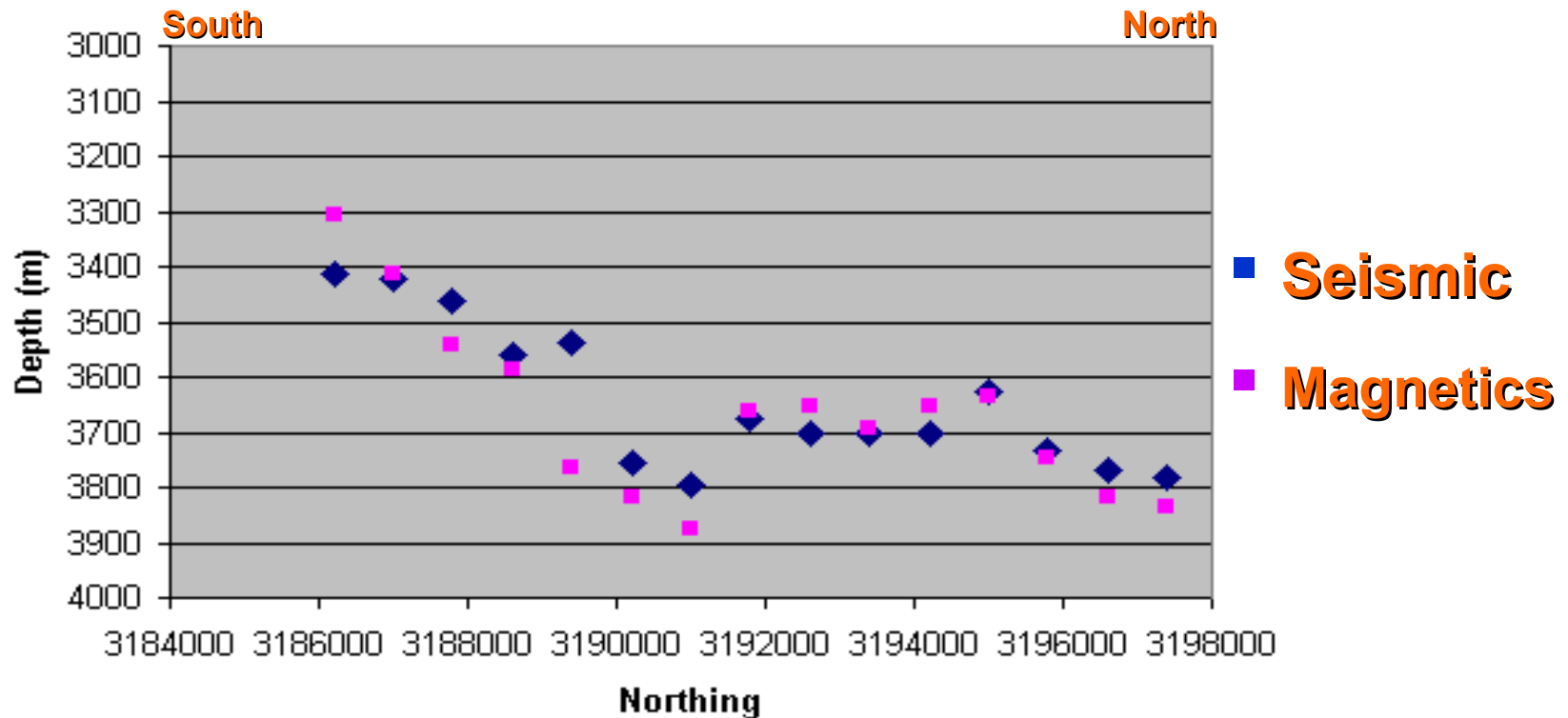
Seismic vs Magnetics

EW PROFILE
Seismic Horizon in Blue
Magnetic Horizon Purple



Top Nubian Seismic vs Magnetics

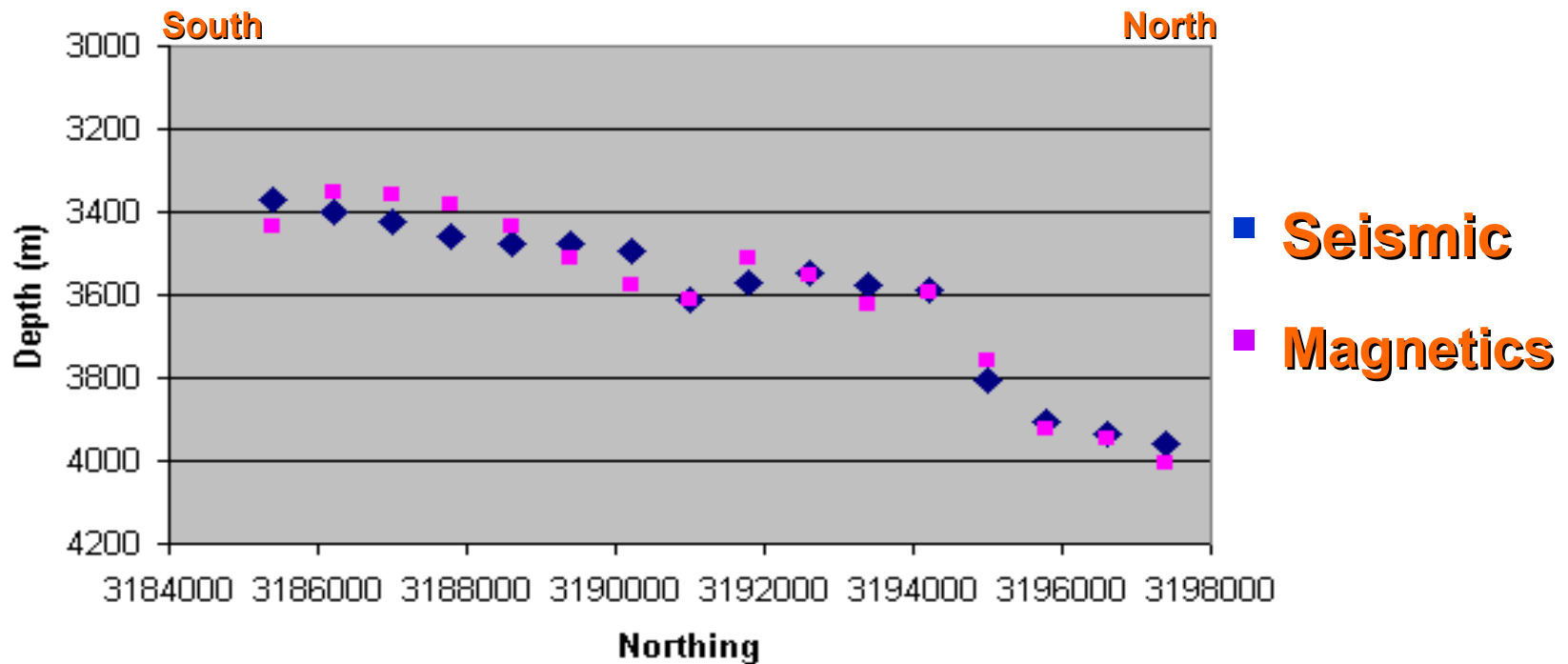
NS PROFILE-1
Seismic Horizon in Blue
Magnetic Horizon Purple



Top Nubian

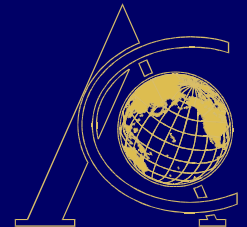
Seismic vs Magnetics

NS PROFILE-2
Seismic Horizon in Blue
Magnetic Horizon Purple



Contents

- Introduction
- Methodology
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Conclusions

- Spectral methods applied to HRAM data can be successfully used in Sirt Basin to image Basement and intra-sedimentary horizons
- Results from ESA Moving Window for Top Nubian SST and Unit 4 are comparable with seismic and well data.
- Multi-Window Test can be used to determine optimal window sizes for spectral analysis.

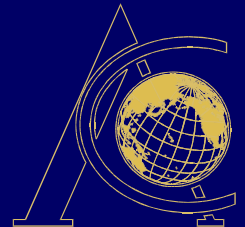


Conclusions

- Automatic spectral interpretation techniques can image multiple horizons:

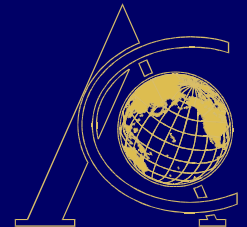
- Basement
- Unit 4 Nubian SST
- Volcanics
- Upper Sabil

- Spectral shape based techniques can identify bodies of limited depth extent such as volcanic layers.

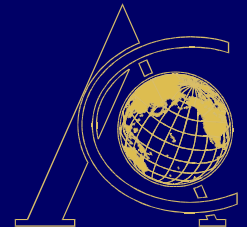


Conclusions

- Energy spectral analysis can be used as a very powerful tool in the analysis and interpretation of HRAM data for exploration in the Libyan Sirt Basin.
- The Multi-Window Test using window density and spectra shape parameters is a new technique which allows the detection in a purely automatic manner the top of Unit 4 of Nubian SST, which is the base of reservoir sands in the study area.



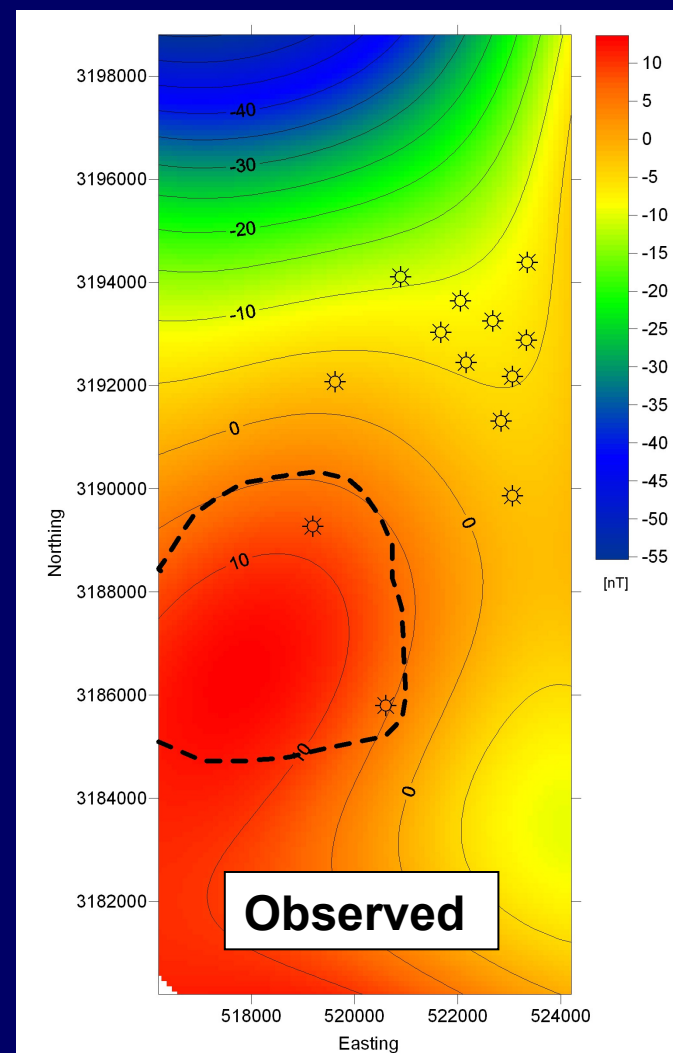
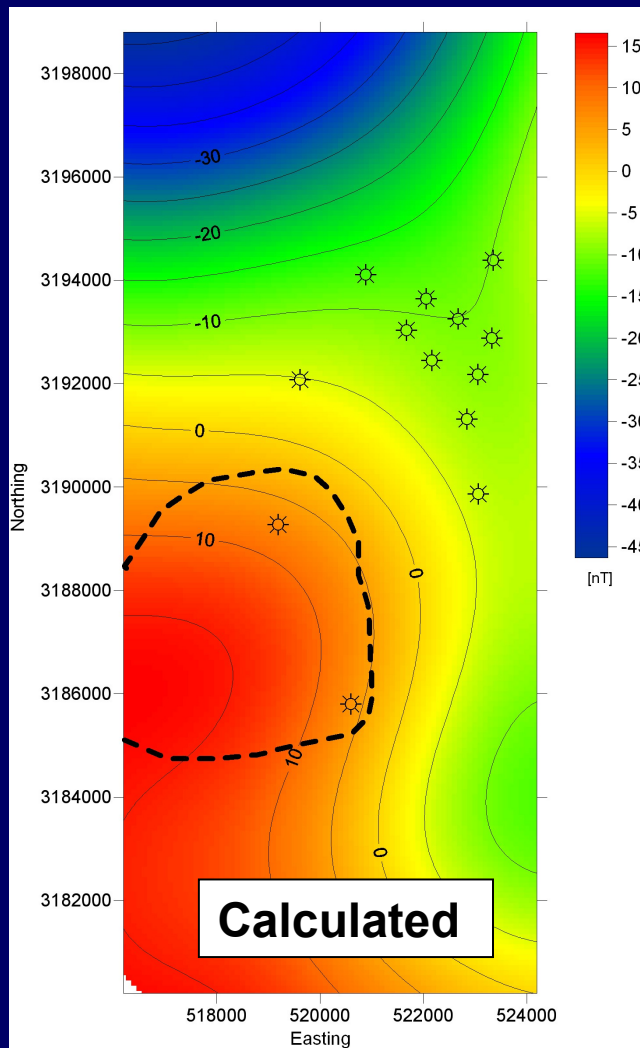
Quality Control



Quality Control: 3D Modelling

Calculated Magnetic Field vs Observed

'Lowpass Filter', Depth > 3260m



--- Interpreted Edge of Volcanics

--- Interpreted Edge of Volcanics



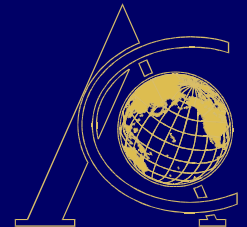
Spectral Ambiguity & Confidence Maps

- Alternative spectra interpretation based on:
 - Spectral shape
 - Consistency with nearby interpretations
 - Geological constraints such as well and seismic data.
- Spectra ambiguity maps quantifying degree to which spectra may admit alternate interpretation



Ambiguity Determination

- Set of alternate semi-automatic depth estimates for 25% spectra
- For each alternate set, interpreter estimates probability that the alternate depth would be chosen over the actual calculated depth
- This data is used to determine a standard deviation [m], that corresponds to the maximum likelihood standard deviation of the normal distribution that would give the actual and alternate depths at the corresponding selected probability ratio.



Spectral Ambiguity Maps

For each Horizon we provide two maps:

- **Spectral ambiguity:**

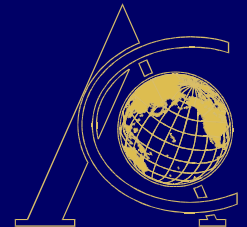
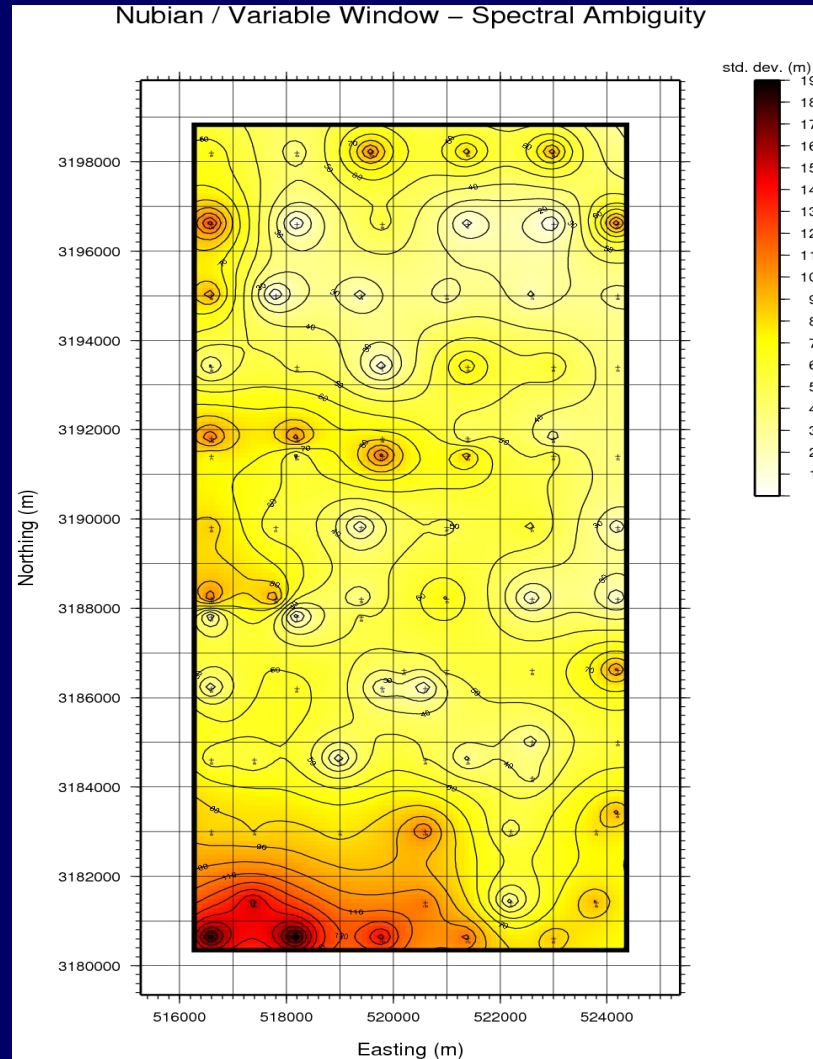
A map of the maximum likelihood estimated standard deviation derived from the actual and alternate semi-automatic depth values. This roughly corresponds to the error-bar on depth estimate that is attributable to spectral ambiguity.

- **Confidence map:**

A map of the probability that a depth distributed according to the maximum likelihood normal distribution will fall within one grid-cell length of the interpreted depth.

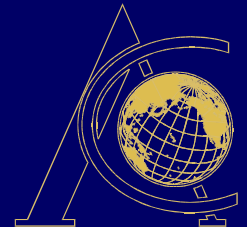
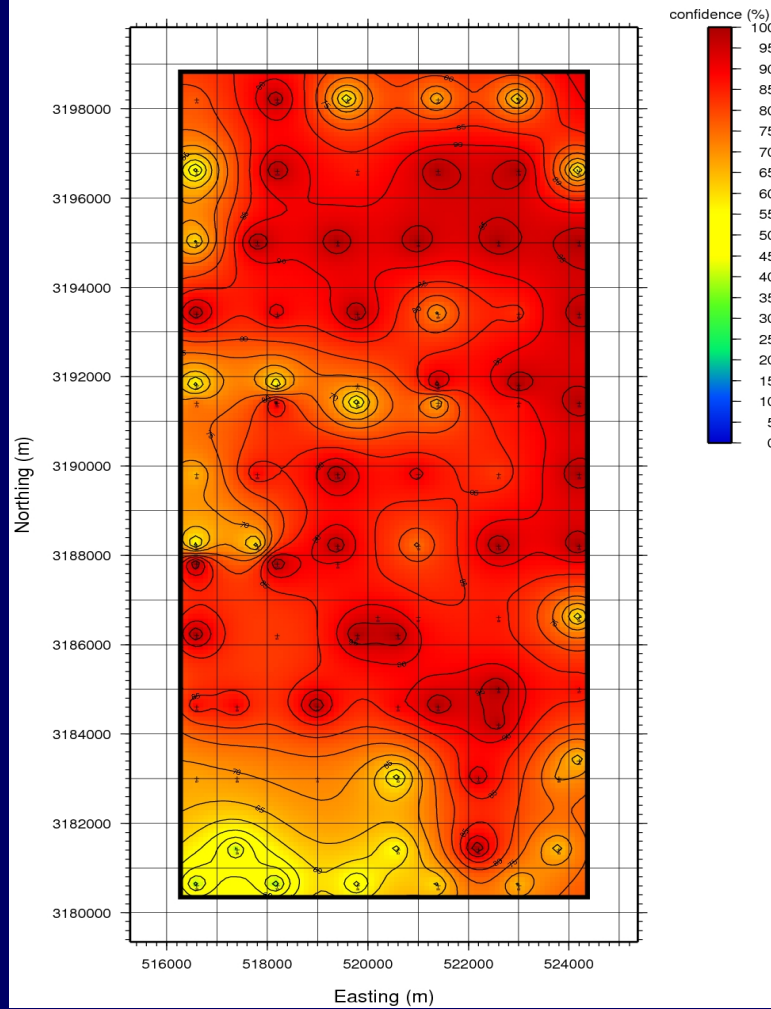


Spectral Ambiguity Map

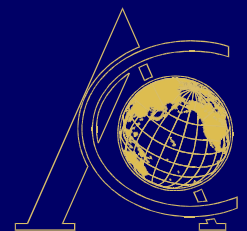
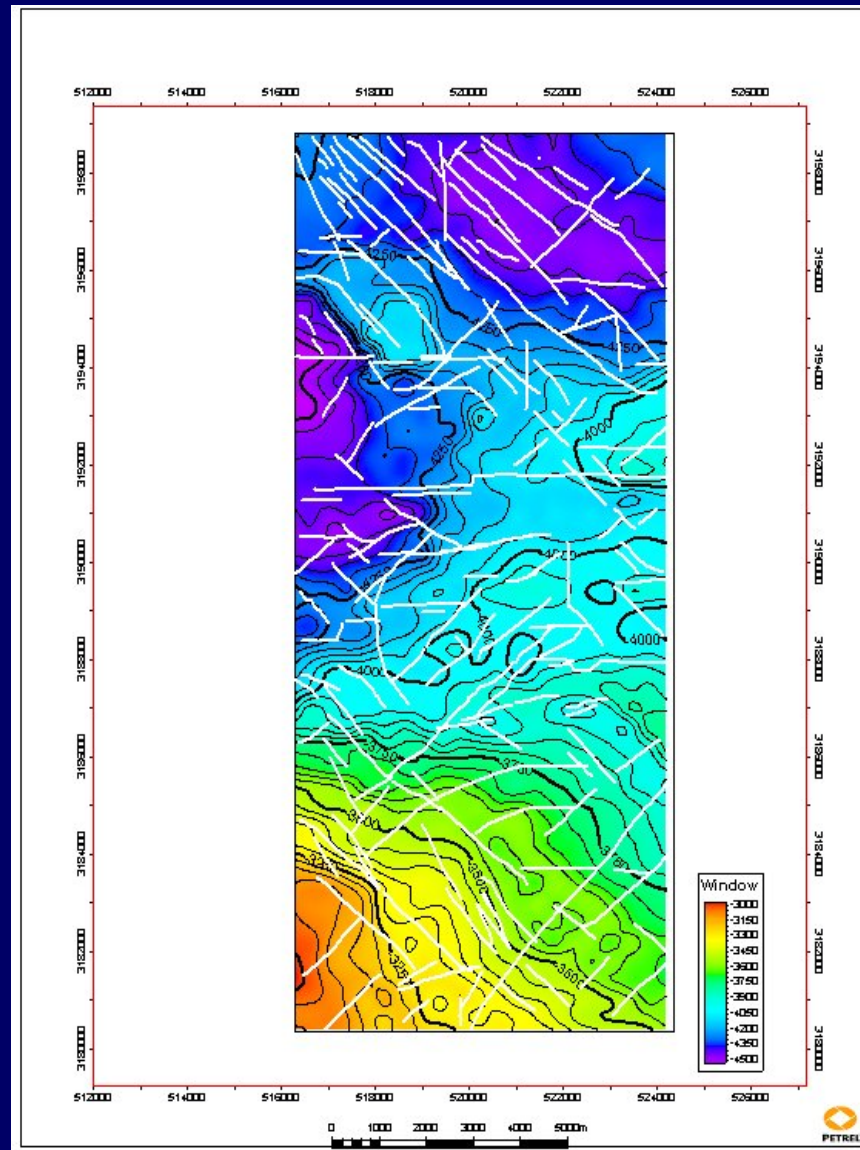


Confidence Map

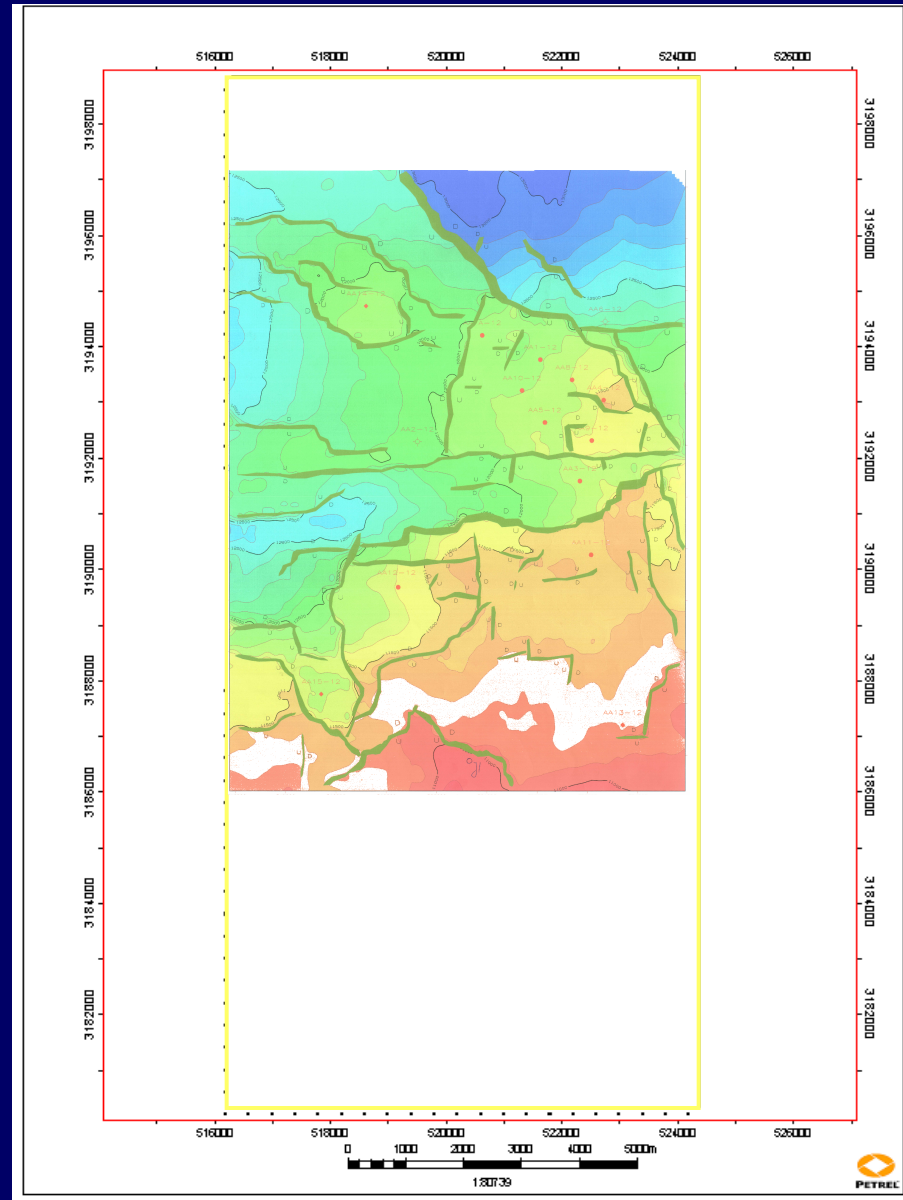
Nubian / Variable Window – Ambiguity 100 m Confidence



Basement plus 1km magnetic lineaments overlay onto BSMT map

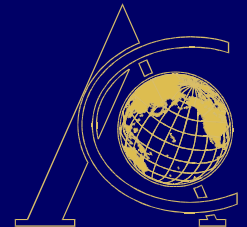


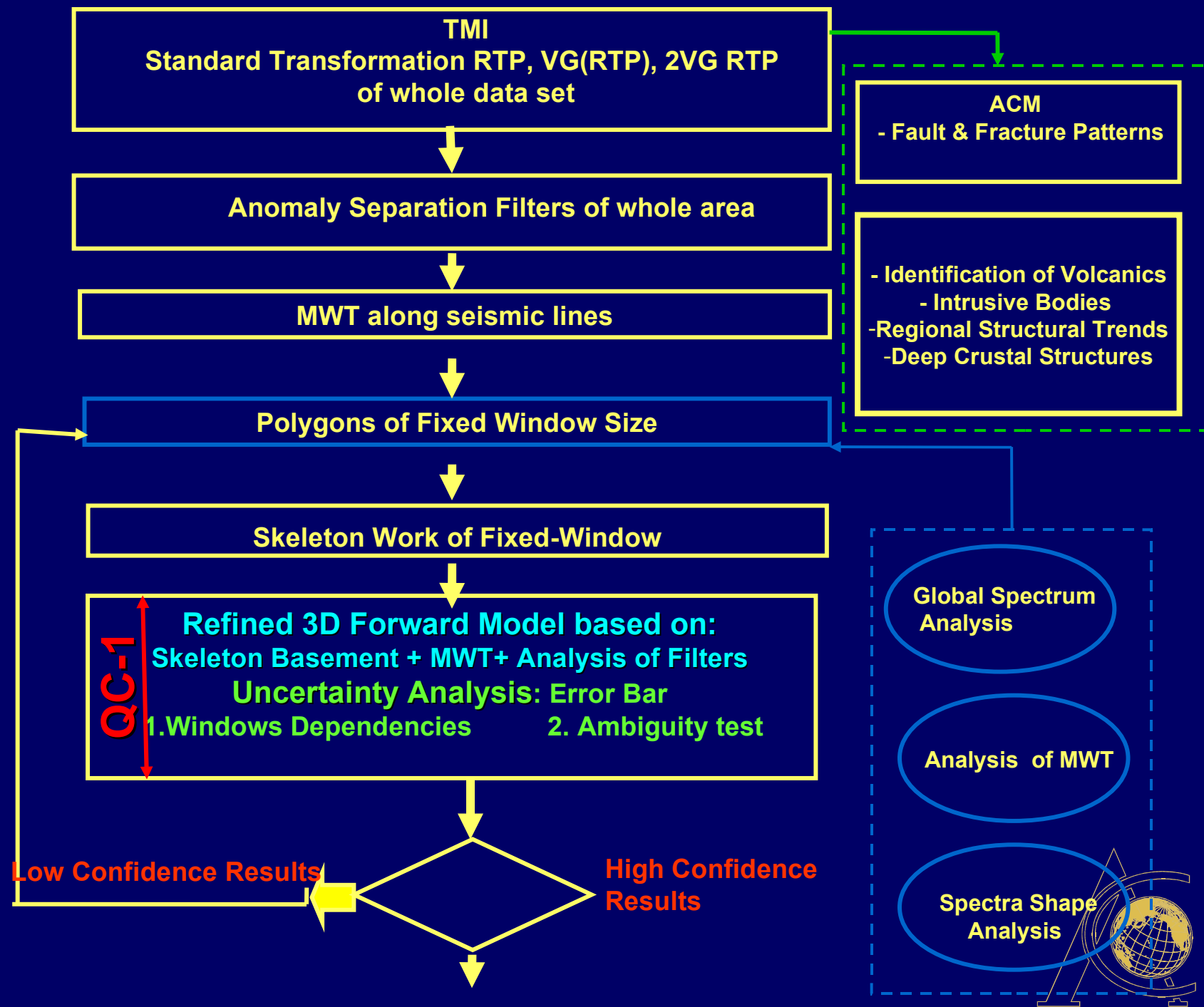
Top Nubian: Seismic



Sedimentary Basin Study

Workflow Chart





Fixed Window Spectra

- Semi-Automatic
- AutoESA

Refined 3D Forward Model based on:
Skeleton of Basement + MWT+ Analysis of Filters Results

Uncertainty Analysis:
Windows Dependencies & Ambiguity test

Comparison with Seismic Data & Well Data

ACM Comparison

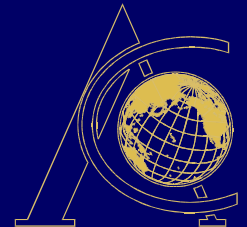
QC-2

Low confidence results

High Confidence Results

Integration Fixed Window Results with Seismic:

- Refining Spectra calculation
- Refining Semi-Automatic Depth Calculation



Analysis of MWT to define Variable Window size



Variable Window Spectra

- Semi-Automatic
- AutoESA



Refined 3D Forward Model based on:

- Skeleton of Basement
- MWT output
- Analysis of Filters Output

Uncertainty Analysis:

- Windows Dependencies
- Ambiguity test

QC-3



Presentation and Report

